

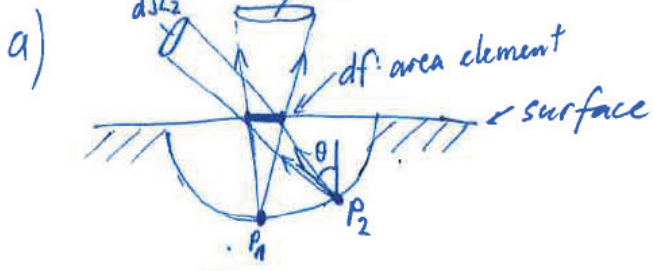
# Homework problem sheet 3

A1

energy density:  $u$ , speed of light (photons):  $c$

solid angle (Raumwinkel) of a sphere:  $4\pi$

$\Rightarrow$  power spectral density of emitted radiation:  $S^*(\nu, T) = \frac{c}{4\pi} u(\nu, T)$



Intuitive: the radiation emitted through a small hole in the surface is proportional to the area of the hole in the direction of the emitted particles.



$$\tilde{df} = df \cdot \cos(\theta)$$

$\Rightarrow$  emitted power per solid angle is  $S^* \cdot \cos(\theta)$

b)  $S = \int_0^\infty d\nu \int_0^{2\pi} d\psi \int_0^{\pi/2} d\theta \cos(\theta) \cdot \frac{c}{4\pi} u(\nu, T) \cdot \sin(\theta)$

Annotations:   
 -  $\int_0^\infty d\nu$ : integrate full spectrum   
 -  $\int_0^{2\pi} d\psi \int_0^{\pi/2} d\theta$ : half a sphere   
 -  $\cos(\theta)$ : from a)   
 -  $\sin(\theta)$ : solid angle element:  $d\Omega = \sin(\theta) \cdot d\theta \cdot d\psi$

$$= \frac{c}{4\pi} \cdot \frac{8\pi h}{c^3} \int_0^{2\pi} d\psi \int_0^{\pi/2} \sin(\theta) \cos(\theta) d\theta \cdot \int_0^\infty \frac{\nu^3}{e^{\frac{h\nu}{kT}} - 1} d\nu = \frac{2\pi h}{c^2} \int_0^\infty \frac{\nu^3}{e^{\frac{h\nu}{kT}} - 1} d\nu$$

$\frac{1}{2} \sin^2 \theta \Big|_0^{\pi/2} = \frac{1}{2}$

$$= \frac{2\pi h}{c^2} \int_0^\infty \frac{\left(\frac{kT}{h}\right)^3 \cdot x^3}{e^x - 1} \cdot \frac{kT}{h} dx = \frac{2\pi h}{c^2} \left(\frac{kT}{h}\right)^4 \int_0^\infty \frac{x^3}{e^x - 1} dx = \frac{2\pi^5 k_B^4}{15 \cdot c^2 h^3} \cdot T^4 \propto T^4$$

$x = \frac{h\nu}{kT}$   
 $\nu = \frac{kT}{h} \cdot x$   
 $d\nu = \frac{kT}{h} \cdot dx$

look up:  $\frac{\pi^4}{15}$

Stefan-Boltzmann constant:  
 $\sigma = 5.67 \cdot 10^{-8} \frac{W}{K^4 m^2}$

c) - Emitted power from the sun:

$$P_{\text{sun}} = S \cdot 4\pi R_{\text{sun}}^2 = 4\pi R_{\text{sun}}^2 \cdot \sigma T_{\text{sun}}^4$$

$\uparrow$  emitted power per area       $\uparrow$  surface area

- power arriving from the sun on the area A at a distance D (earth):

$$P_{\text{earth}} = A \cdot \frac{P_{\text{sun}}}{4\pi D^2} = \sigma T_{\text{sun}}^4 \cdot \left(\frac{R_{\text{sun}}}{D}\right)^2 \cdot A$$

- losses in the conversion to electricity:

$$\underline{P_{\text{el}}} = P_{\text{earth}} \cdot \eta_{\text{atm}} \cdot \underbrace{\eta_{\text{conv}} \cdot \eta_{\text{BW}}}_{\text{"efficiency" of solar cell} \approx 20\%} = \underline{1.39 \text{ MW}}$$

d) Power emitted from the sun:  $P_0 = S \cdot A_0 = \sigma \cdot T_0^4 \cdot 4\pi R_0^2$

$\uparrow$  emitted power/area       $\uparrow$  surface of the sun       $\uparrow$  Stefan-Boltzmann

Fraction of light entering the eye:  $\tilde{P}_{\text{eye}} = P_0 \cdot \frac{A_{\text{eye}}}{4\pi D^2} = P_0 \cdot \frac{(d/2)^2 \pi}{4\pi D^2}$

Factor to account for atmosphere, spectral efficiency of the eye...:  $\eta = 0.9$

$\Rightarrow$  power entering the eye:  $P_{\text{eye}} = \eta \cdot \sigma T^4 \cdot 4\pi R_0^2 \cdot \frac{d^2 \pi}{4\pi D^2}$

$\Rightarrow$  number of photons:  $\underline{r} = P_{\text{eye}} / h\nu_{\text{max}} = P_{\text{eye}} \frac{\lambda_{\text{max}}}{hc} = \frac{\lambda_{\text{max}}}{4hc} \eta \sigma T^4 \frac{R_0^2 d^2 \pi}{D^2} \approx 4.38 \cdot 10^{15} \text{ photons/s}$

e)

$$P_{\text{max}} = 1 \frac{\text{mW}}{\text{cm}^2} = 10 \frac{\text{W}}{\text{m}^2}$$

$\Rightarrow r_{\text{max}} = \underbrace{P_{\text{max}} / h\nu}_{\text{photon rate per area}} \cdot \underbrace{(d/2)^2 \pi}_{\text{area}} \approx 7.90 \cdot 10^{13} \text{ photons/s} \ll r_{(d)}$

photon rate per area      area

$\Rightarrow$  it is dangerous to stare into the sun (ask Galileo!)

A2

$$a) E_i = \frac{hc}{\lambda} = 17.45 \text{ keV}$$

$$b) \text{ complete backscattering: } \theta = \pi \text{ (cos } \theta = -1) \implies \underline{\lambda_f} = \lambda_i + \frac{2h}{mc} = \underline{\underline{0.07595 \text{ nm}}}$$

$$c) E_f = \frac{hc}{\lambda_f} = 16.34 \text{ keV}$$

$$d) p = \hbar k = \hbar \cdot \frac{2\pi}{\lambda} = \frac{h}{\lambda}$$

$$p_i = \frac{h}{\lambda_i} = 9.31 \cdot 10^{-24} \frac{\text{kg} \cdot \text{m}}{\text{s}}$$

$$p_f = \frac{h}{\lambda_f} = 8.72 \cdot 10^{-24} \frac{\text{kg} \cdot \text{m}}{\text{s}} \text{ (in opposite direction)}$$