

Dye sensitized solar cells (2)

Basic Principles of Photovoltaics

Maximum power conversion of photovoltaic devices

Ressources:

Peter Würfel, Physik der Solarzellen

Jenny Nelson, The Physics of Solar Cells

Thomas Dittrich, lecture notes

Energy demand

of a human

100 W to sustain biological life
(2.4 kWh/day or 876 kWh/year)

>3 kW to sustain comfortable life
(>72 kWh/day or >26.28 MWh/year)

of mankind

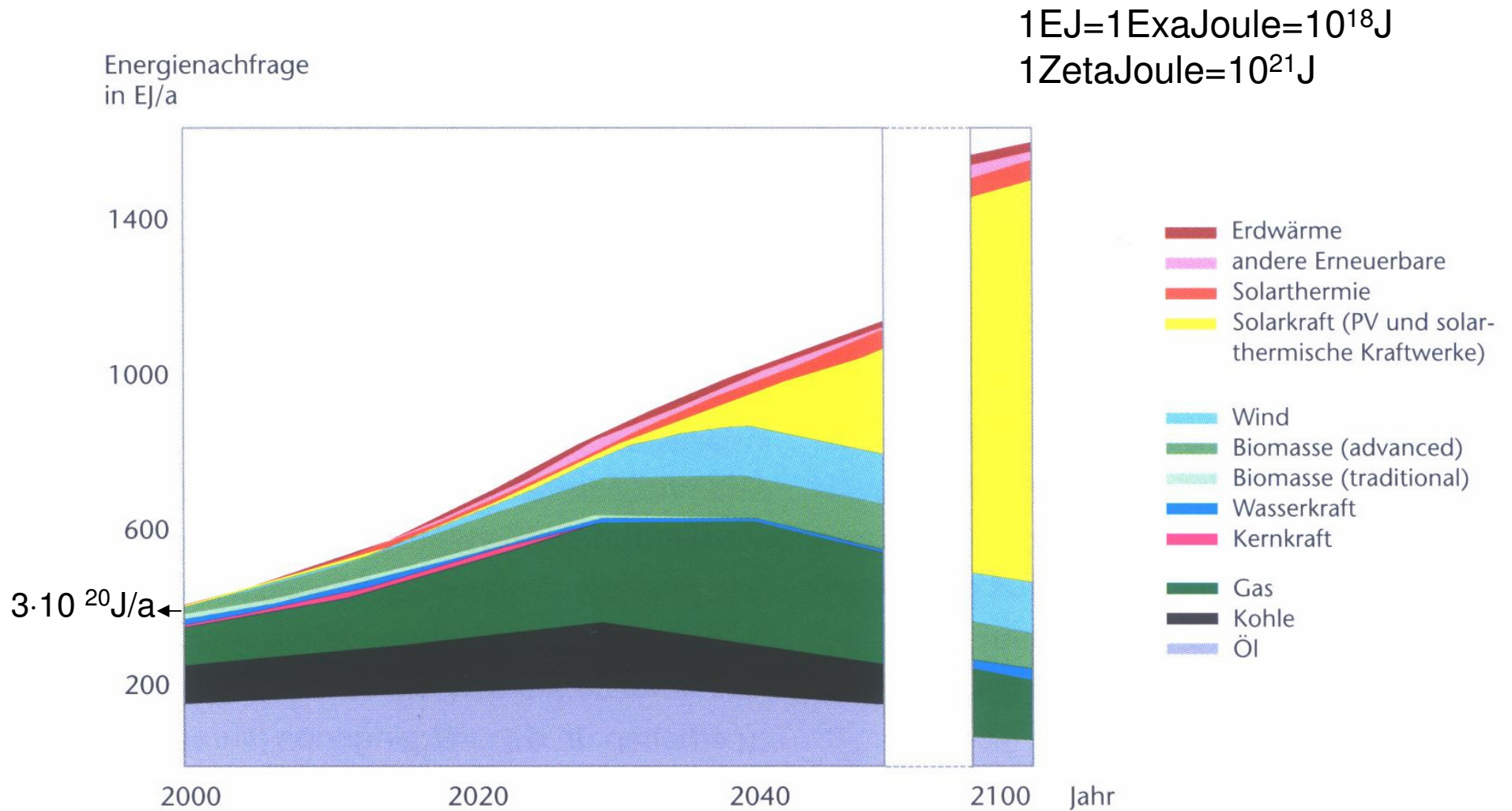
more than 6×10^9 individuals

> 1.8×10^4 GW to sustain comfortable life
(18TW)

for comparison

energy in 100 g chocolate about 1 kWh
energy in 1 l gasoline about 10 kWh
power of one nuclear power plant about 1 GW

Scenario of worldwide energy demand



300 ExaJoule per year correspond about to 9 TW = 9000GW

The sun is a hot fire ball

spectrum of the black body radiation (Planck)
(energy flux as a function of the photon energy and of the temperature)

$$\frac{dJ_S(h\nu)}{d(h\nu)} = \frac{8\pi}{h^3 c^2} \cdot \frac{(h\nu)^3}{\exp\left(\frac{h\nu}{k_B T_S}\right) - 1}$$

$$(h\nu)_{\max} = 2.82 \cdot k_B T$$

with

$$h = 6.6 \cdot 10^{-34} \text{ Js}$$

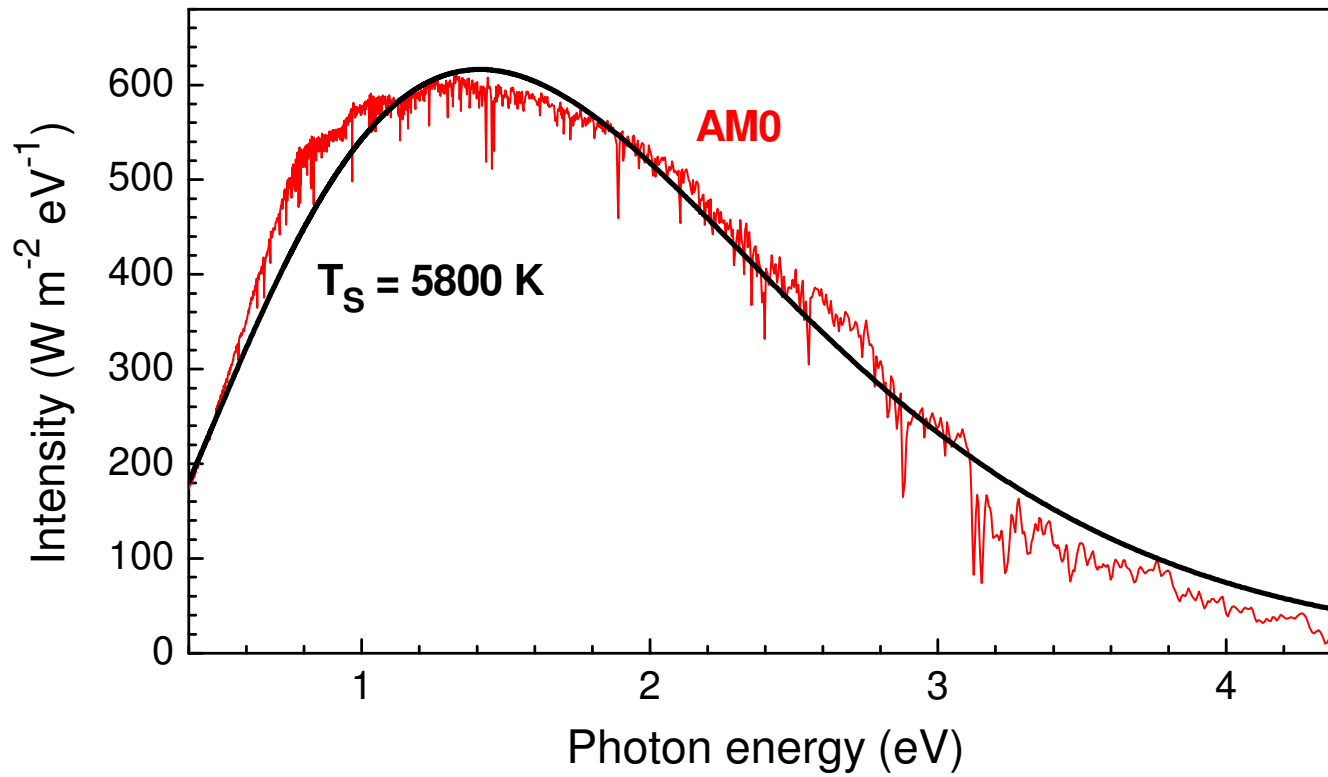
$$c = 3 \cdot 10^8 \text{ m/s}$$

$$k_B = 1.38 \cdot 10^{-23} \text{ J/K}$$

By measuring the photon energy at the maximum of the sun spectrum T_s can be obtained.

temperature at the surface of the sun:

$$T_s \approx 5800 \text{ K}$$



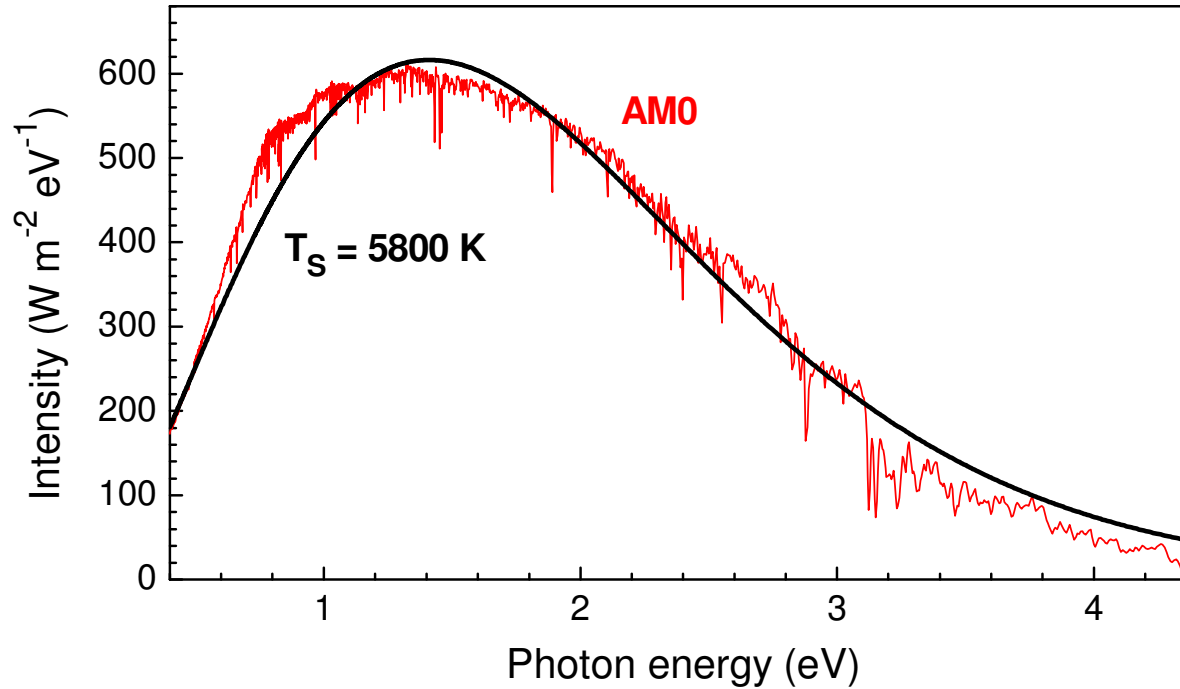
The unit of the photon flux

$$[d\Phi_S] = \frac{[dJ_S]}{[h\nu]} = \frac{W}{m^2 \cdot eV} \cdot \frac{1}{[e]V}$$

$$[d\Phi_S] = \frac{AV}{m^2 \cdot eV} \cdot \frac{1}{AsV}$$

$$[d\Phi_S] = \frac{1}{cm^2 \cdot eV \cdot s} = \frac{\text{number of photons}}{\text{unit area} \cdot \text{energy unit} \cdot \text{time unit}}$$

Energy flux spectrum and photon flux



for example:

$$dJ_s(1.2 \text{ eV}) \approx 550 \cdot \frac{\text{W}}{\text{m}^2 \cdot \text{eV}}$$

$$b_s(1.2 \text{ eV}) \approx \frac{550}{10^4 \cdot 1.2 \cdot 1.6 \cdot 10^{-19} \cdot \text{cm}^2 \cdot \text{eV} \cdot \text{s}}$$

$$b_s(1.2 \text{ eV}) \approx 3 \cdot 10^{17} \cdot \frac{1}{\text{cm}^2 \cdot \text{eV} \cdot \text{s}}$$

Power of the sun received on earth

Stefan-Boltzmann law $I_S = \sigma \cdot T^4$

power of the sun $P_S = \sigma \cdot T^4 \cdot 4\pi \cdot R_s^2$

power received on earth $P_e = P_S \cdot \frac{\pi \cdot R_e^2}{4\pi \cdot (AU)^2}$

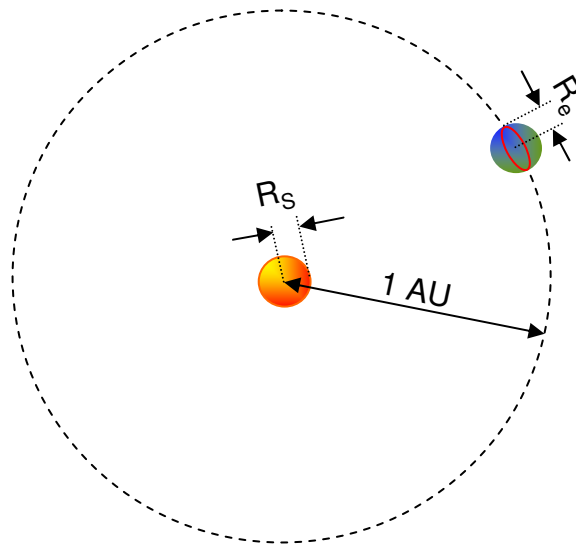
with

$\sigma = 5.67 \cdot 10^{-8} \text{ W}/(\text{m}^2\text{K}^4)$

$R_s = 7 \cdot 10^5 \text{ km}$

$R_e = 6.4 \cdot 10^3 \text{ km}$

$AU = 1.5 \cdot 10^8 \text{ km}$



$P_e \approx 1.8 \cdot 10^8 \text{ GW}$

more than 1000...10000 times
the energy demand of mankind

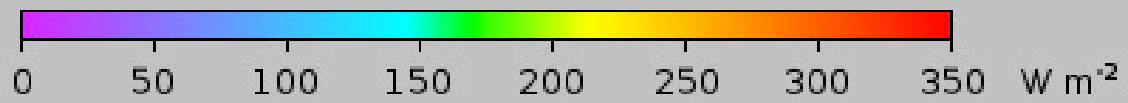
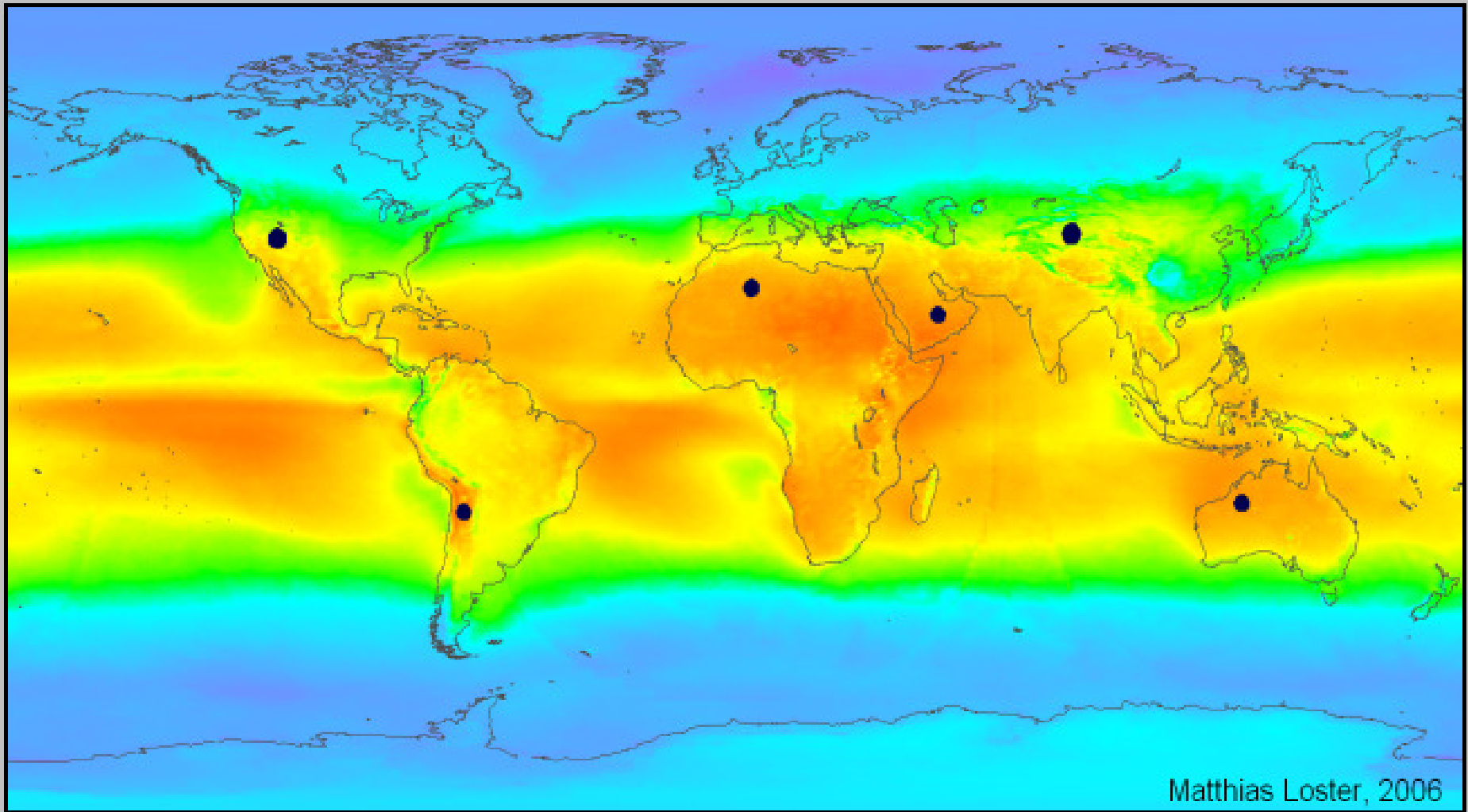
At present mankind needs about 9000GW

The solar constant

energy flux of sun light to the earth normalized to the area of 1 m²

$$J_s = \frac{P_s}{4\pi \cdot (AU)^2}$$

$$\mathbf{J_s = 1.353 \text{ kW/m}^2}$$



$\Sigma \bullet = 18 \text{ TWe}$

Annual average solar irradiance

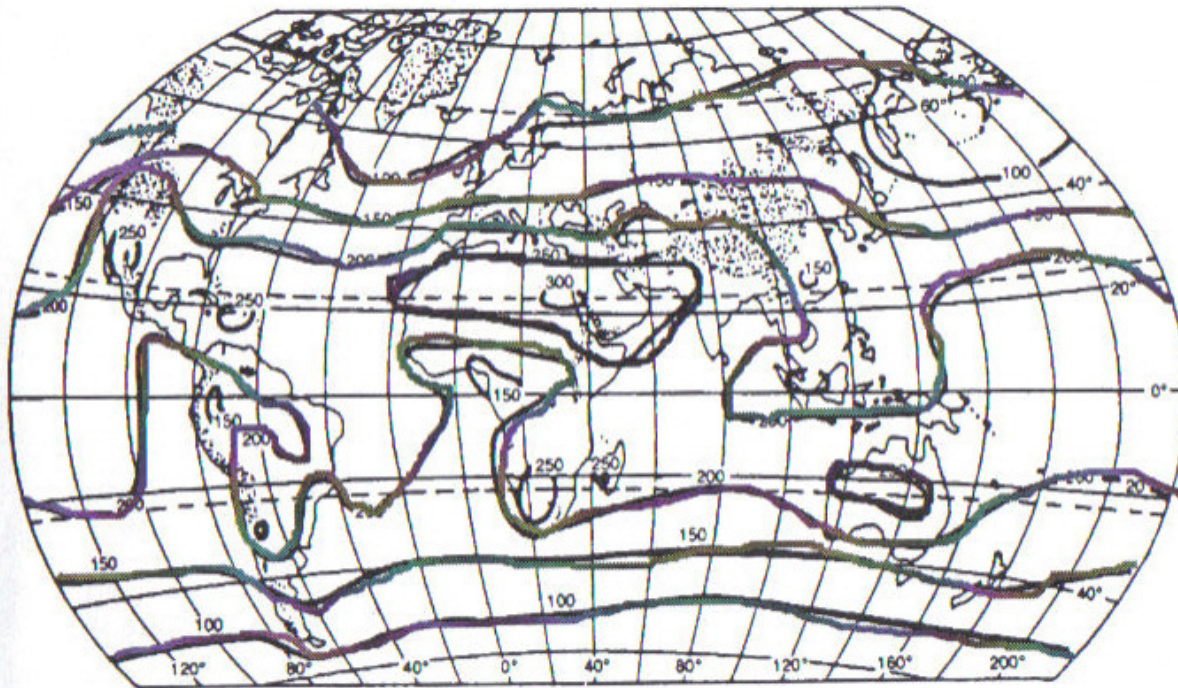


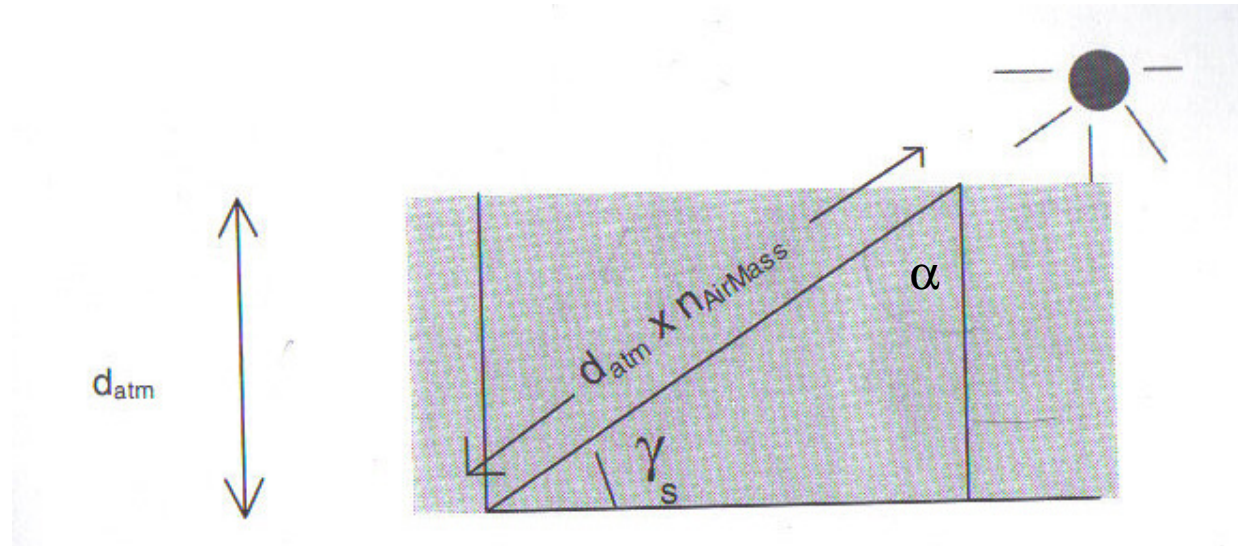
Fig. 2.3. Global distribution of annual average solar irradiance. The values on the irradiance contours are given in W m^{-2} .

Global average: Extraterrestrial $\langle J_E \rangle = 1/4 \cdot 1353 \text{W/m}^2 = 338 \text{W/m}^2$

CH: 115W/m^2

Saudi-Arabien: 285W/m^2

Air Mass: AM1.5, AM1.0....



$$n_{\text{air mass}} = 1 / \sin(\gamma_s)$$

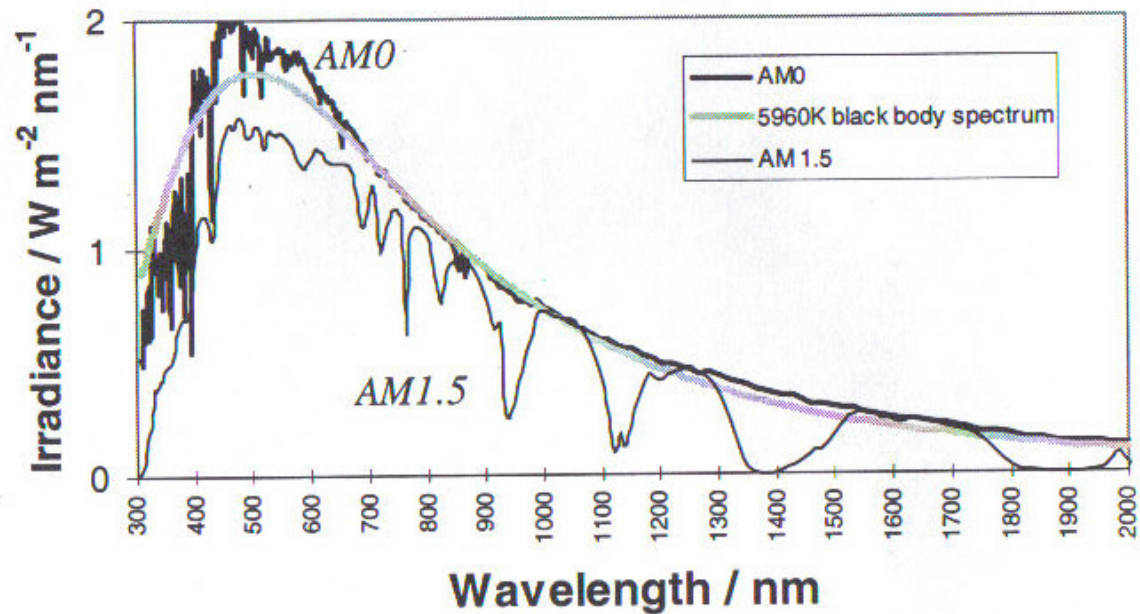
$$\gamma_s = 41.8^\circ \Rightarrow n_{\text{air mass}} = 1.5$$

$$\text{Alternatively: } n_{\text{air mass}} = 1 / \cos(\alpha)$$

Definition: For convenience AM1.5 is defined as 1000W/m² integrated irradiance
(Typical experimental values at 42° are 900W/m²)

AM0: Extraterrestrial; AM1.0: equator

Solar spectrum



- Comparison of AM0 (extraterrestrial) with black body radiation of $T=5960\text{K}$ and AM 1.5
- Dips in spectrum at 900, 1100, 1400 and 1900nm (H_2O) and at 1800 and 2600nm (CO_2)

Temperature on earth from energy balance

$$I_{\text{atm}} \rightarrow 0$$

Energy flux from sun to earth

$$I_s = J_s \cdot \pi \cdot R_e^2$$

Black body radiation from earth into free space

$$I_e = \sigma \cdot T_e^4 \cdot 4\pi \cdot R_e^2$$

with

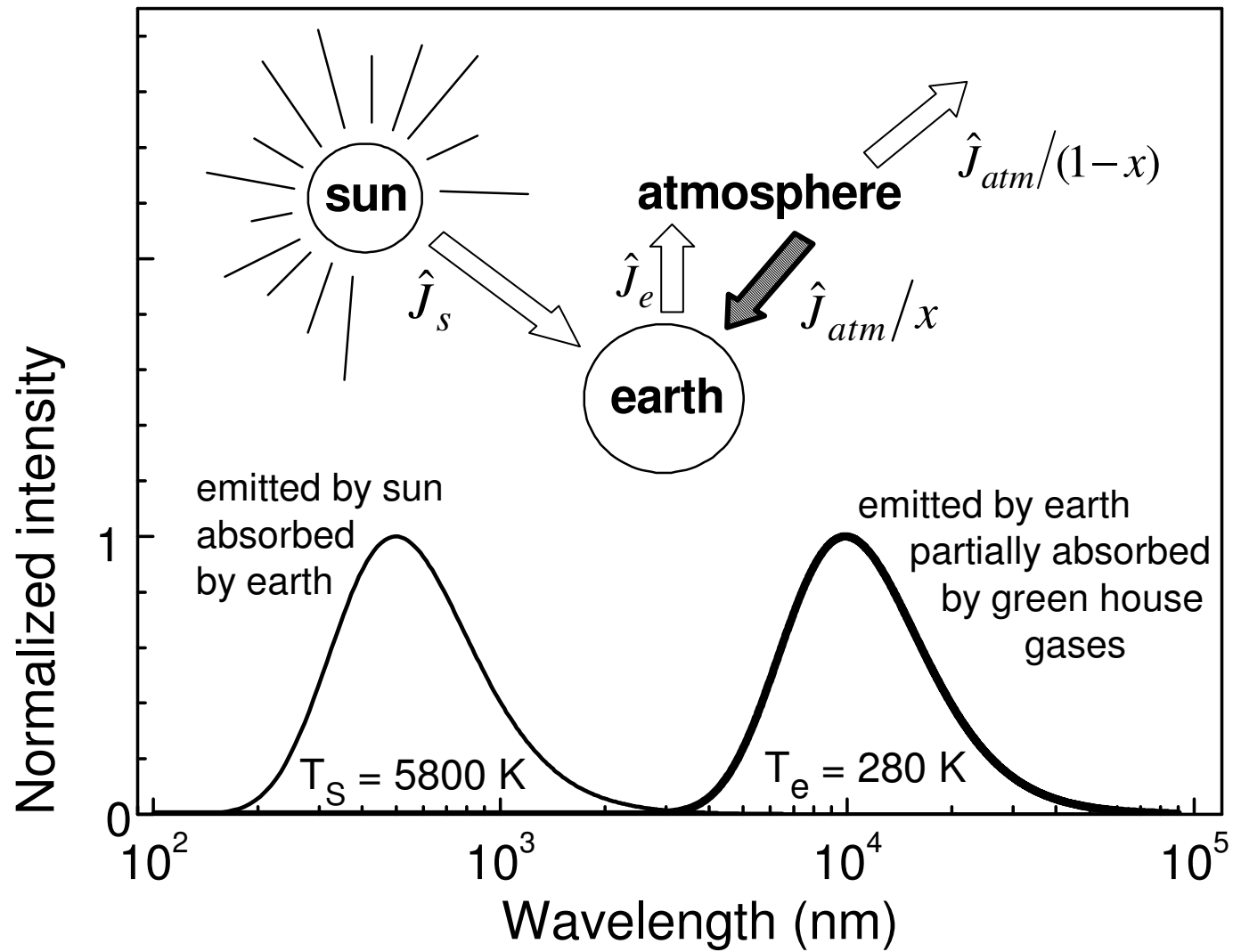
$$\sigma = 5.67 \cdot 10^{-8} \text{ W}/(\text{m}^2\text{K}^4)$$

$$J_s = 1.35 \text{ kW}/\text{m}^2$$

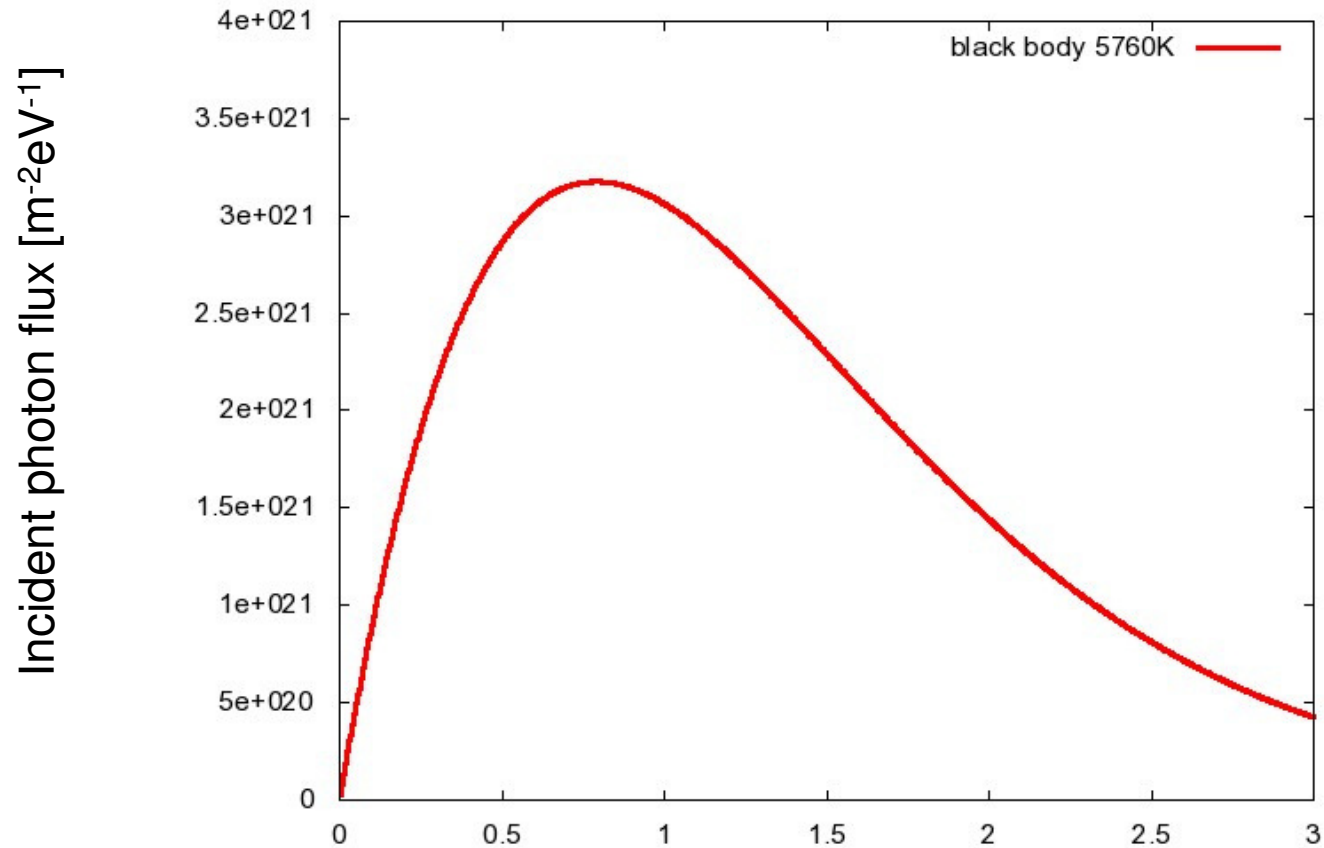
Balance

$$I_e = I_s$$

$$\Rightarrow T_e = \sqrt[4]{\frac{J_s \cdot \pi \cdot R_e^2}{\sigma \cdot 4\pi \cdot R_e^2}} \approx 277 \text{ K}$$



Black body radiation at 5760K



$$b_s(E) = \frac{2F_s}{h^3 c^2} \left(\frac{E^2}{e^{E/k_B T_s} - 1} \right) \quad E(\text{eV})$$

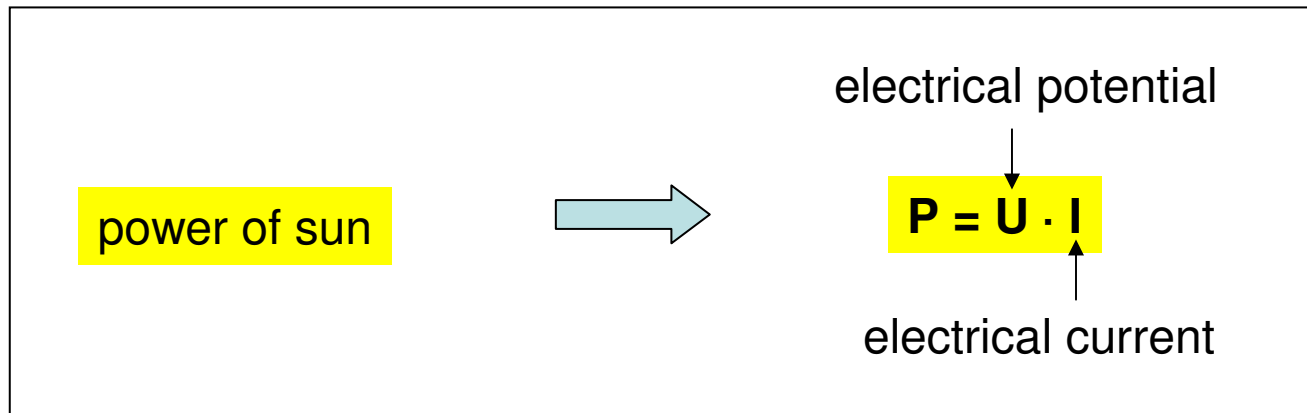
$$F_s = 6.8 \cdot 10^{-5}$$

(corresponds to the sun-earth geometry)

*

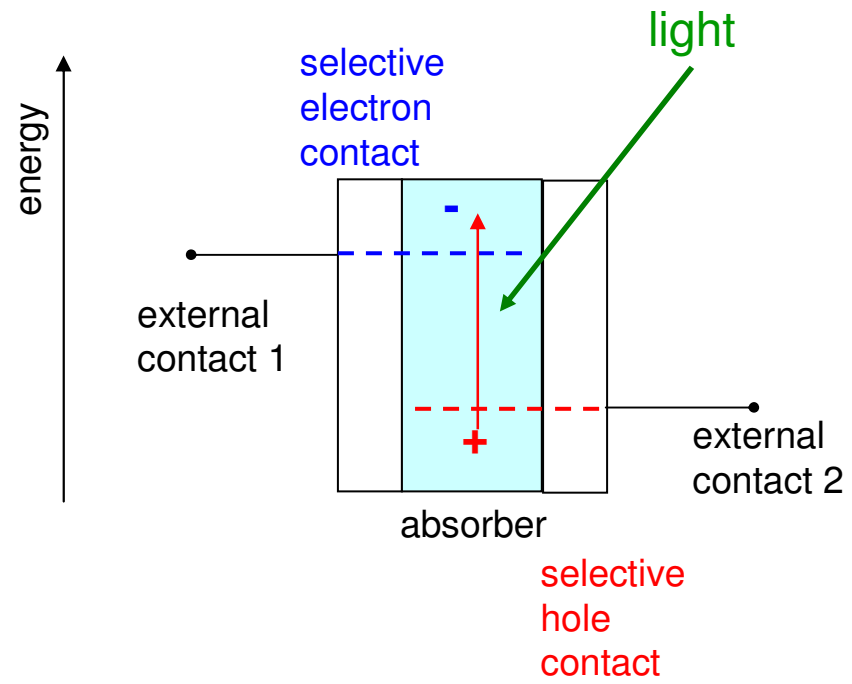
The solar cell

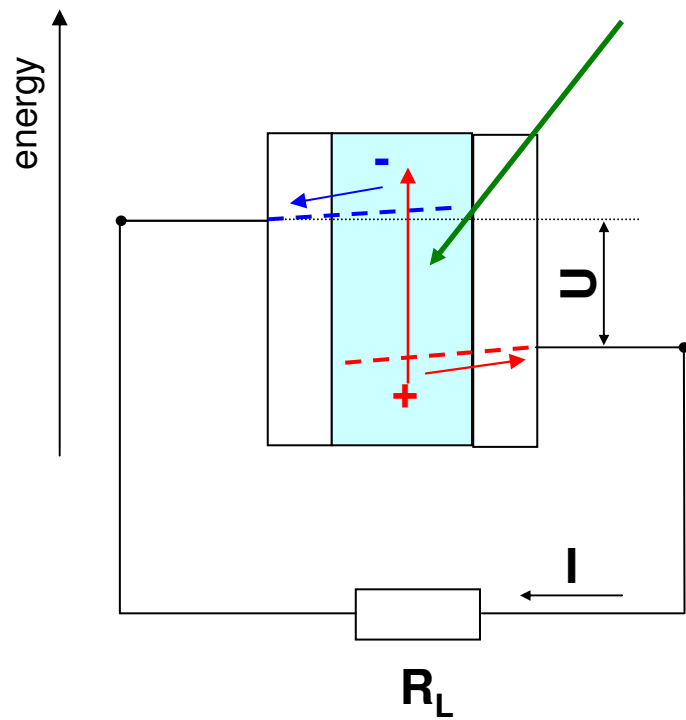
electronic device converting solar power directly into electrical power



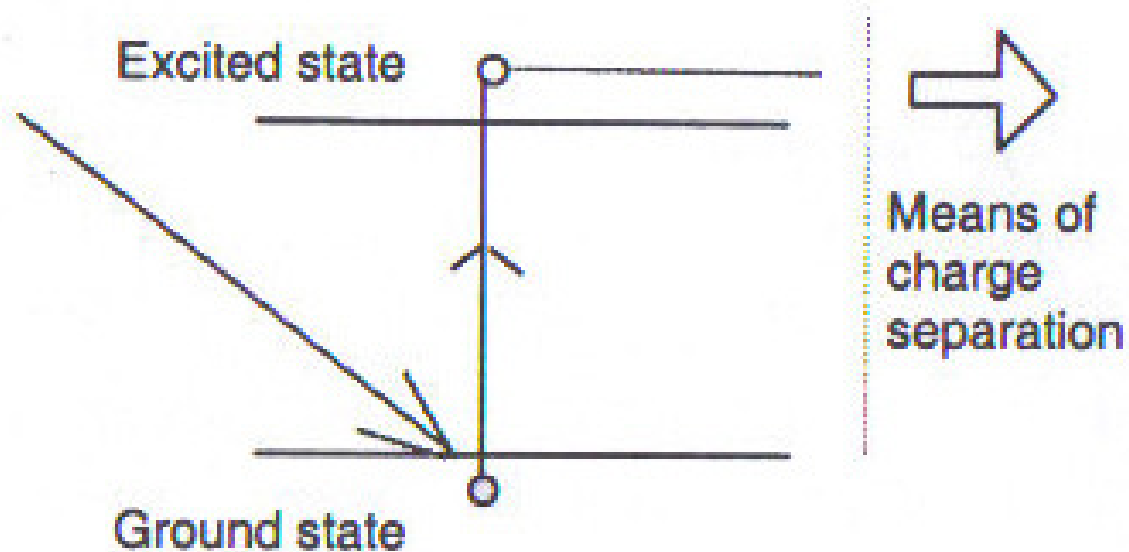
A solar cell absorbs light and conducts electrons and holes to separate electron and hole collecting contact.

A major characteristic is the energy conversion efficiency.





Excitation and charge separation



Photon creates electron-hole pair, which is separated by an asymmetric potential. Charge separation prevents relaxation to its initial state.

Cell in the dark: Absorption of thermal photons

$$b_a(E) = \frac{2F_a}{h^3 c^2} \left(\frac{E^2}{E^{E/k_B T_a} - 1} \right)$$

b_a : incident photon flux

h : Planck's constant

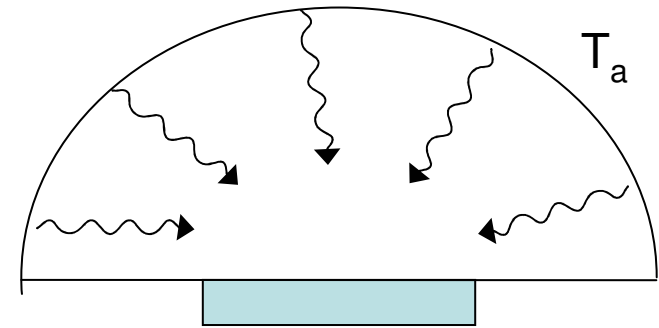
c : speed of light

k_B : Boltzmann constant

E : Energy

Thermal photons are absorbed from the environment of Temperature T_a

F_a : geometry factor; $F_a = \pi$ for hemisphere



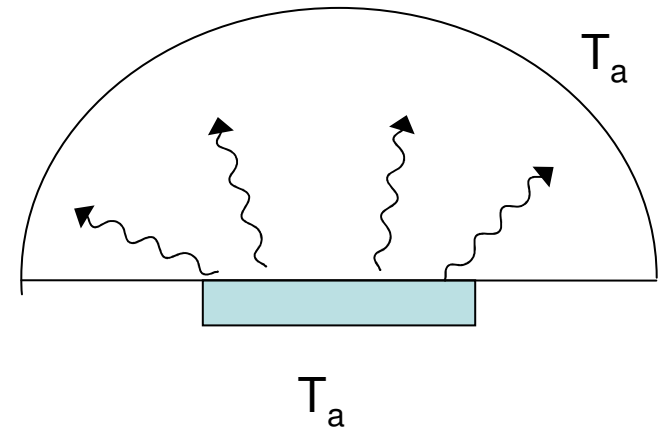
The equivalent current density j_{abs} absorbed from the ambient:

$$j_{abs}(E) = q(1 - R(E))a(E) \frac{2F_a}{h^3 c^2} \left(\frac{E^2}{E^{E/k_B T_a} - 1} \right)$$

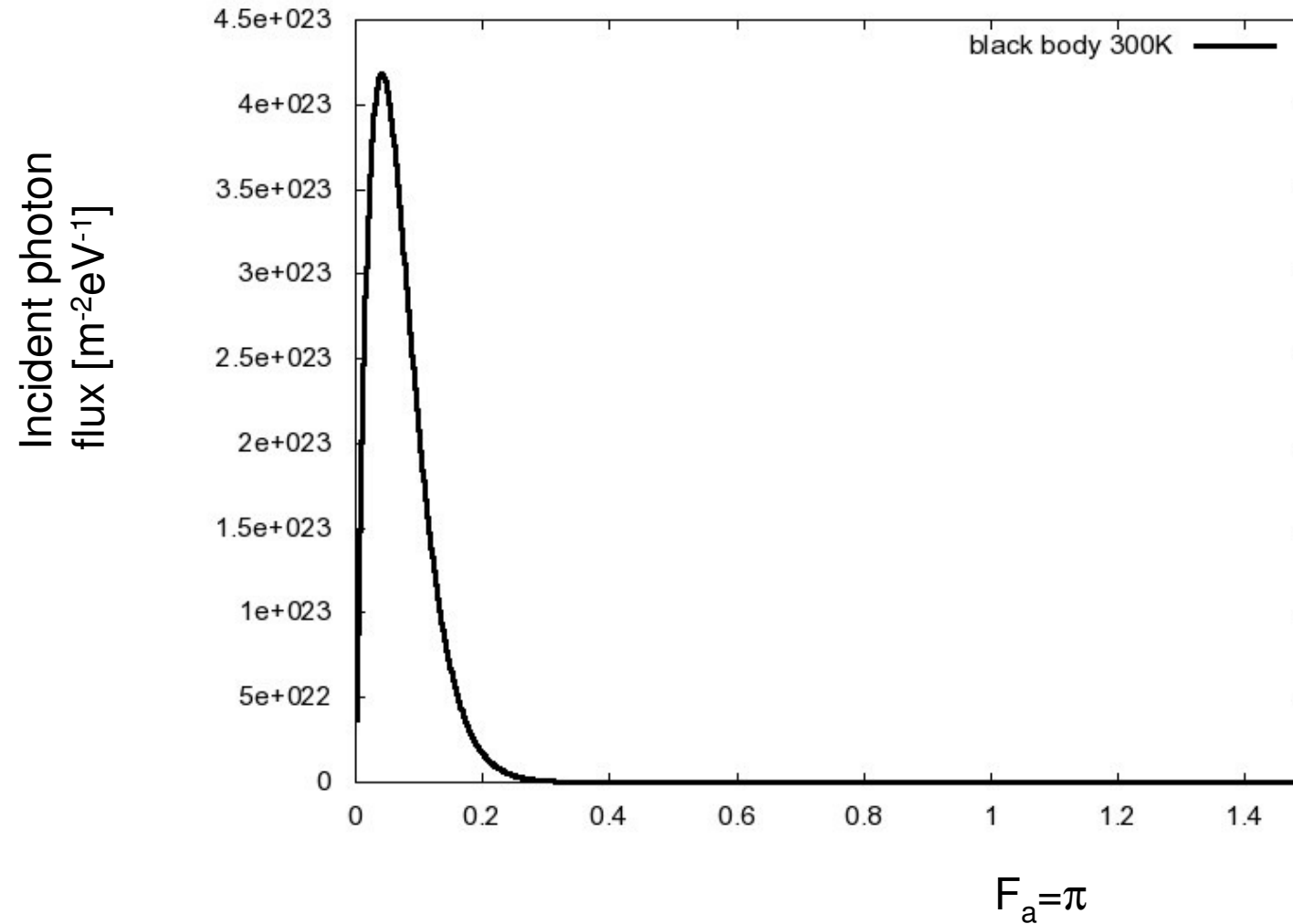
Cell in the dark: Emission of thermal photons

$$j_{rad}(E) = q(1 - R(E))\epsilon(E) \frac{2F_a}{h^3 c^2} \left(\frac{E^2}{E^{E/k_B T_a} - 1} \right) = q(1 - R(E))\epsilon(E)b_a(E)$$

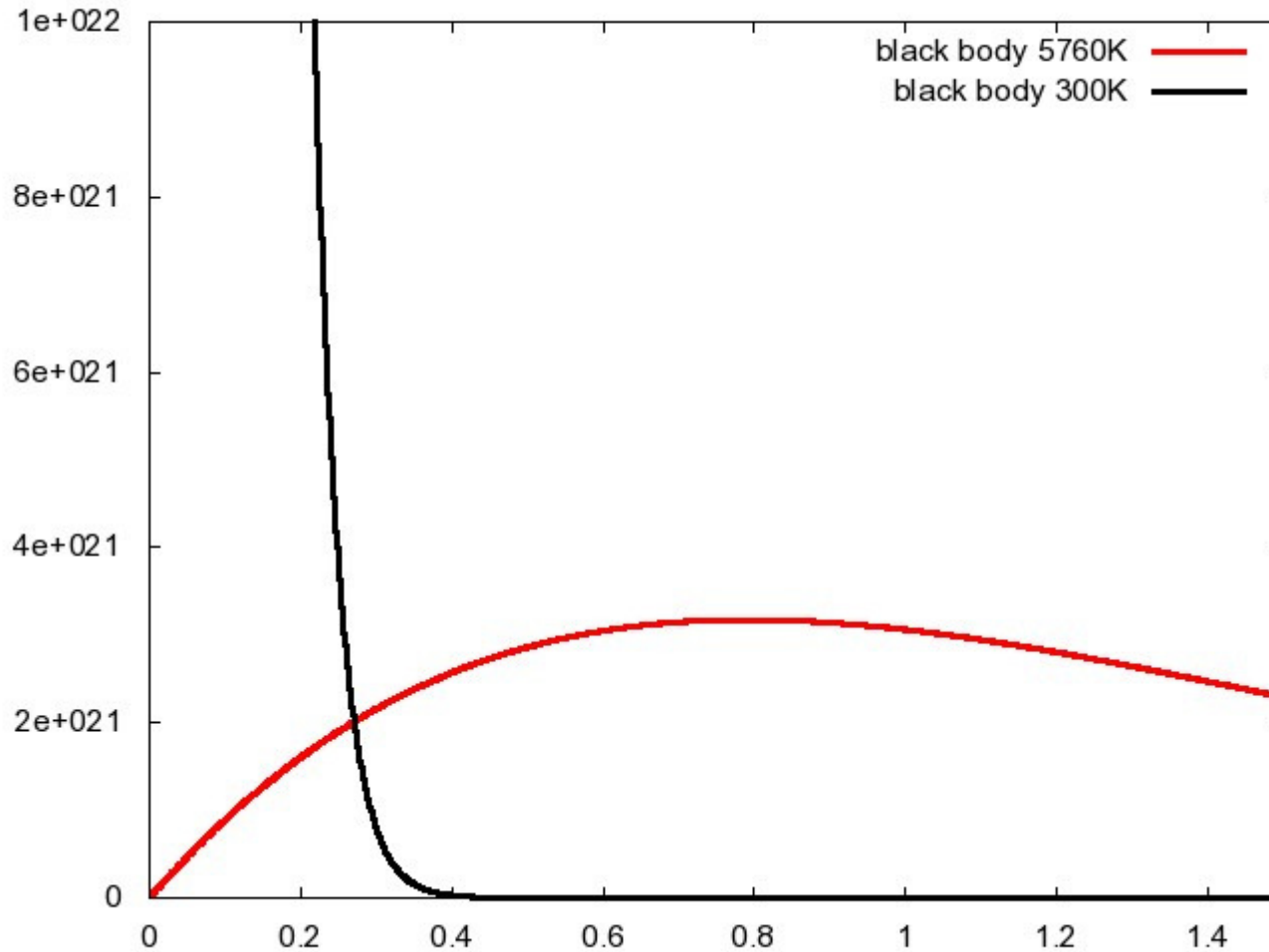
In equilibrium: Kirchoff's law $\epsilon(E)=a(E)$



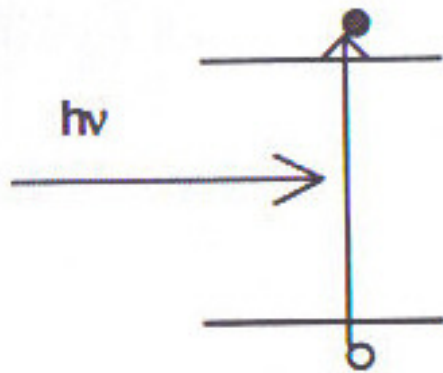
Black body radiation at 300K



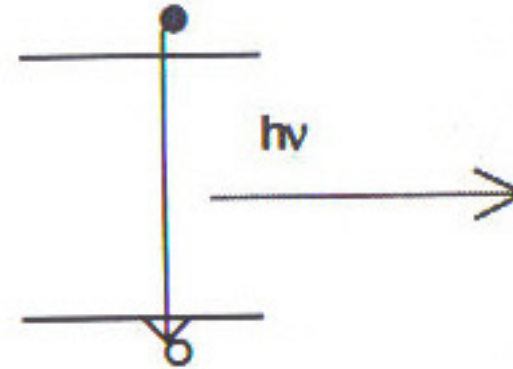
Comparison of black body radiation at 300K and 5760K



Absorption and spontaneous emission



absorption



spontaneous emission

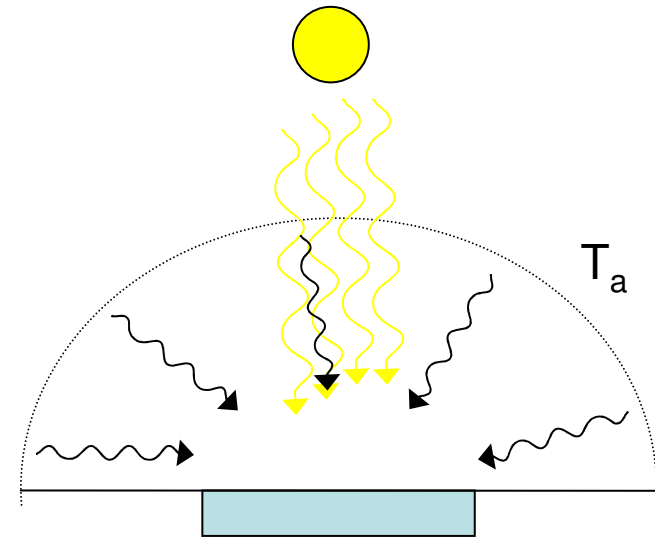
Under illumination (absorption)

$$b_s(E) = \frac{2F_s}{h^3 c^2} \left(\frac{E^2}{E^{E/k_B T_s} - 1} \right)$$

T_s : Temperature of the sun

$$F_s = \pi \sin^2 \theta_s = \pi / 4.6 \cdot 10^4$$

$$\theta_s = 0.26^\circ$$



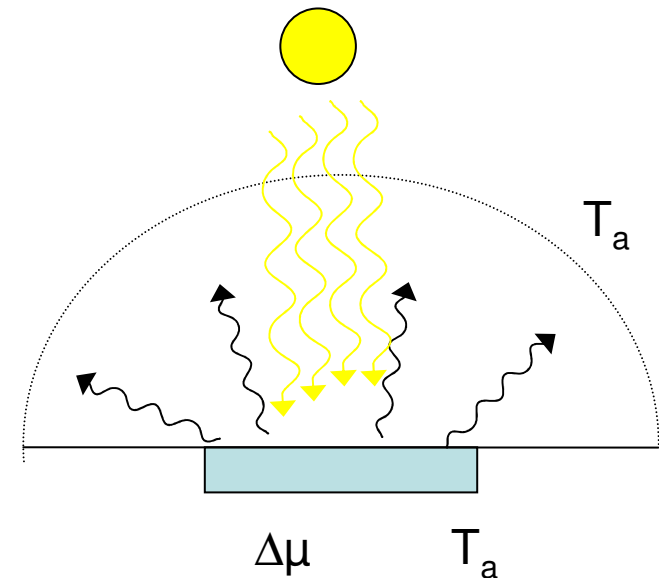
$$j_{abs}(E) = q(1 - R(E))a(E)(b_s(E) + (1 - F_s / F_a)b_a(E))$$

Under illumination (emission)

Depends on the $\Delta\mu$. When electrons are at raised energy, relaxation is more frequent: Generalized Planck's radiation law:

$$b_e(E) = \frac{2F_a}{h^3 c^2} \left(\frac{E^2}{E^{(E-\Delta\mu)/k_B T_a} - 1} \right)$$

$$j_{rad}(E) = q(1 - R(E))\epsilon(E)b_e(E, \Delta\mu)$$



Net current density under illumination

Under the assumption that $\epsilon(E)=a(E)$ and $\Delta\mu=\text{const}$ across the cell

$$j_{abs}(E) - j_{rad}(E) = q(1 - R(E))a(E)(b_s(E) + (1 - F_s / F_a)b_a(E) - b_e(E, \Delta\mu))$$

Net absorption:

$$j_{abs-net}(E) = q(1 - R(E))a(E)(b_s(E) - \frac{F_s}{F_a}b_a(E))$$

Net Emission
Radiative recombination

$$j_{rad-emission}(E) = q(1 - R(E))a(E)(b_e(E, \Delta\mu) - b_e(E, 0))$$

Photocurrent

$$J_{sc} = q \int_0^{\infty} \eta_c(E)(1-R(E))a(E)(b_s(E) - \frac{F_s}{F_a}b_a(E))dE \approx q \int_0^{\infty} \eta_c(E)(1-R(E))a(E)(b_s(E))dE$$

η_c : probability that electron is collected

cf.

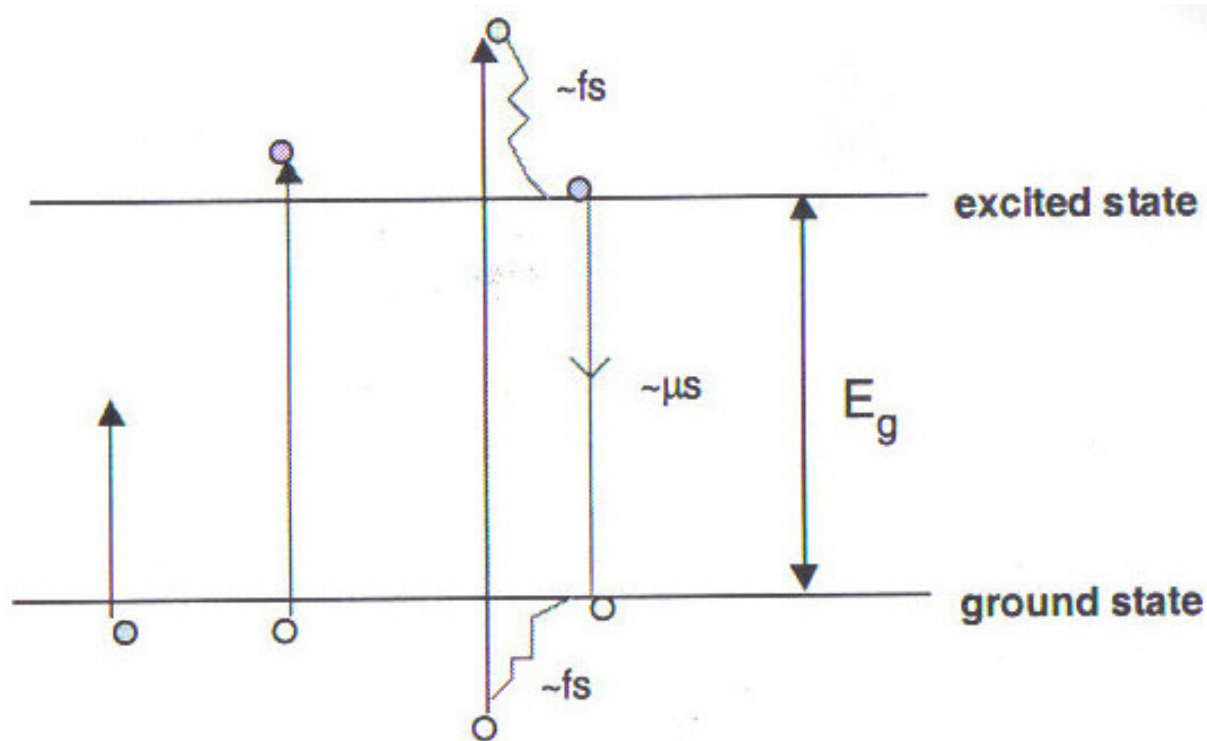
$$J = q \int_0^{\infty} QE(E) \cdot b_s(E) dE$$

Quantum efficiency: $QE(E) = \eta_c(E)(1-R(E))a(E)$

For non-reflecting material: $R(E)=0$ and perfect charge separation $\eta_c=1$:

$$QE(E) = a(E) = \begin{cases} 1 & E \geq E_g \\ 0 & E < E_g \end{cases}$$

Two band photoconverter



Photons with $E \gg E_g$ have the same effect like electrons with $E = E_g$
Photocurrent is rather proportional to photon flux and not photon energy
But photons with $E < E_g$ can not create electron hole pairs

Dark current

Assumption:

-only loss process is radiative relaxation of electrons through spontaneous emission

-No non-radiative losses due to scattering at defects in the material

$$J_{dark} = q \int_0^{\infty} (1 - R(E)) a(E) (b_e(E, \Delta\mu) - b_e(E, 0)) dE$$

where $\Delta\mu = qV$

Dark current is related to emitted flux, which is increased with increasing voltage

Net cell current density

$$J(V) = J_{sc} - J_{dark}(V) = q \int_0^{\infty} (1 - R(E)) a(E) (b_s(E) - (b_e(E, \Delta\mu) - b_e(E, 0))) dE$$

↑
↑
 incoming flux emitted flux

For step-like absorption ($E > E_g$):

$$J(V) = J_{sc} - J_{dark}(V) = q \int_{E_g}^{\infty} b_s(E) - (b_e(E, \Delta\mu) - b_e(E, 0)) dE$$

$J(V)$ depends strongly on voltage:

$$J(V) = J_{sc} - J_0 \left(e^{qV/k_B T} - 1 \right)$$

at $V = V_{oc}$ the total emitted flux balances the absorbed flux

Limiting efficiency

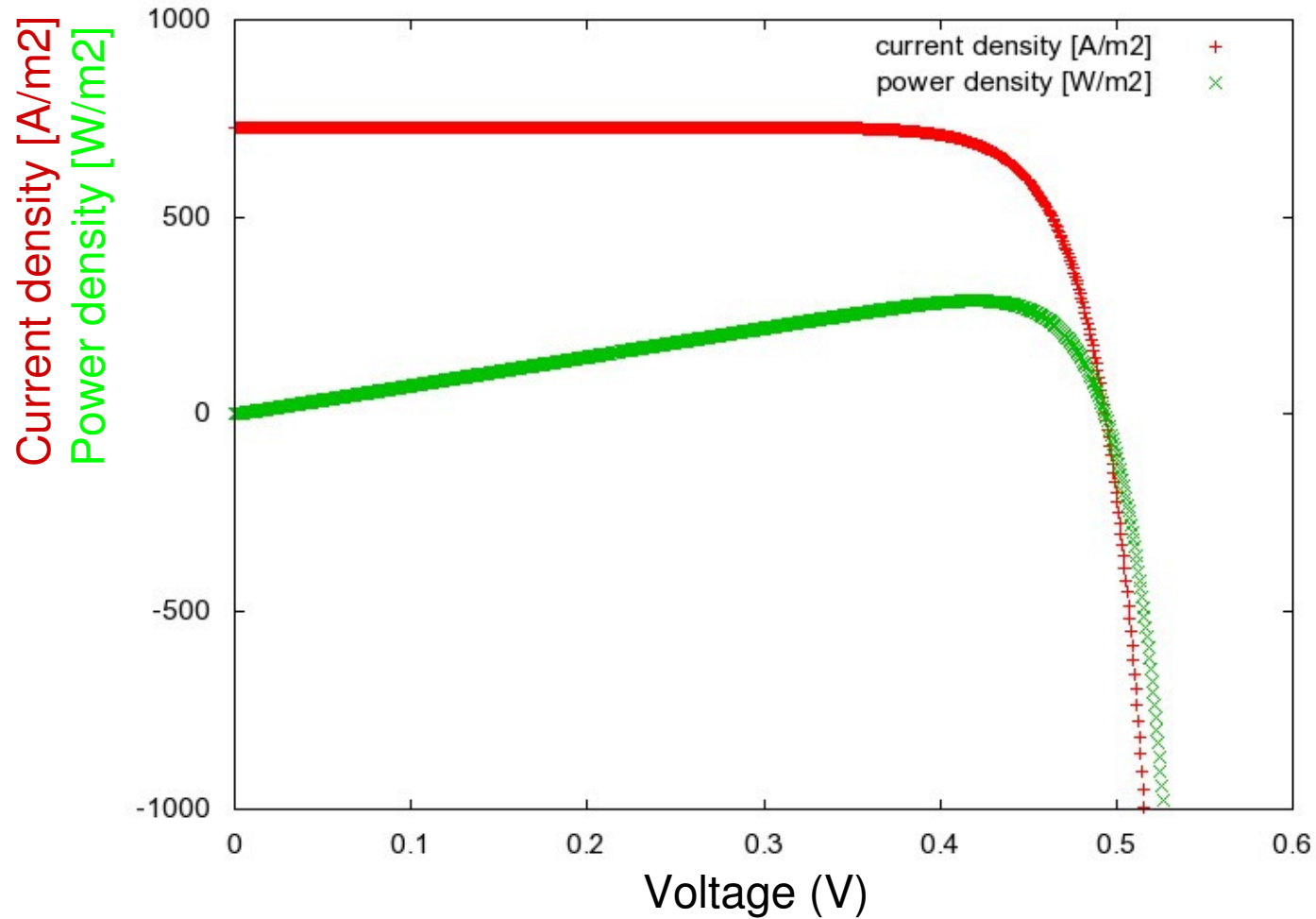
Incident power: $P_s = \int_0^{\infty} E \cdot b_s(E) dE$

Output power: $P = V J(V)$
(assumption no Ohm's losses)

Power conversion efficiency: $\eta = \frac{VJ(V)}{P_s}$

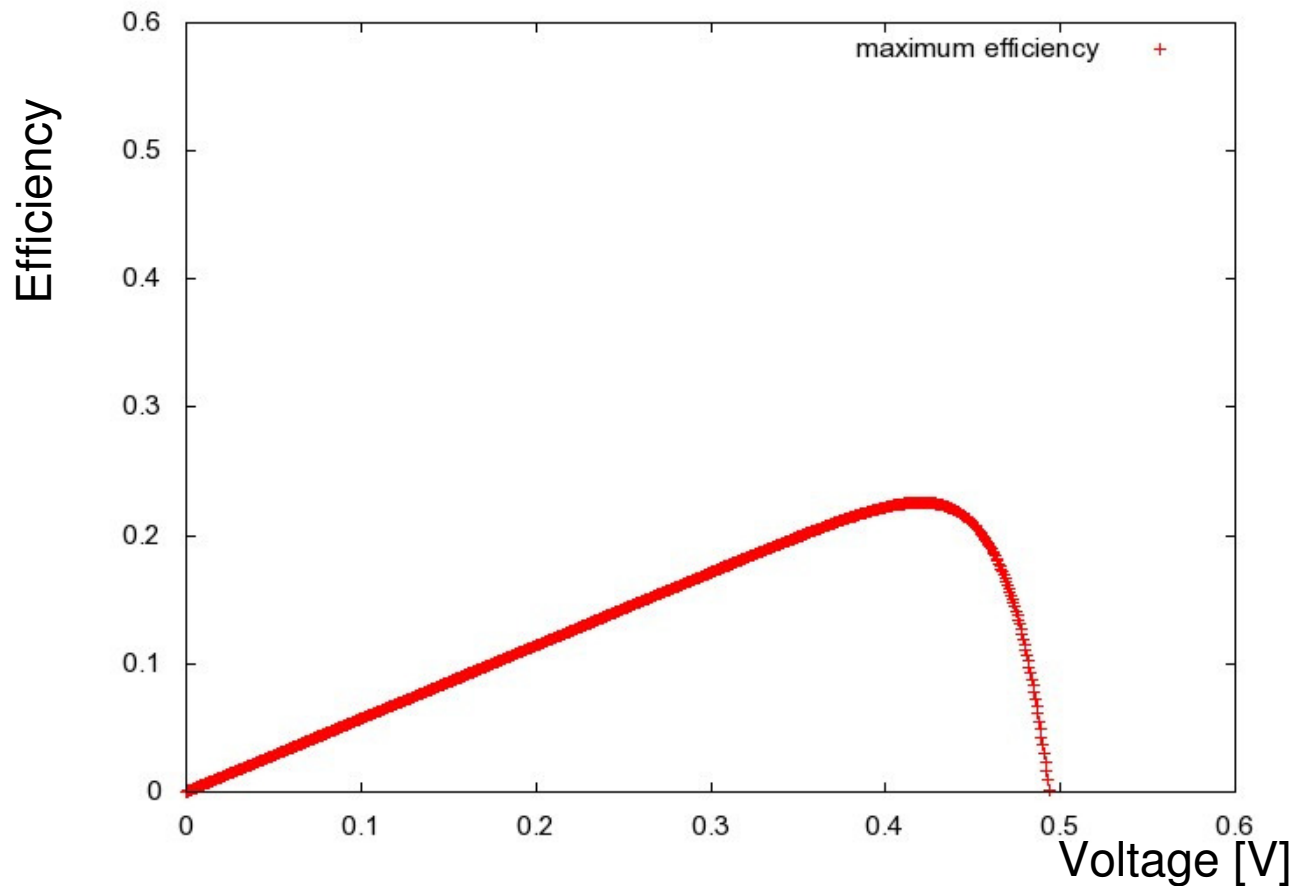
Maximum efficiency is achieved when: $\frac{d}{dV} (J(V)V) = 0$

Maximum current and power densities with $E_g=0.7\text{eV}$



$E_g=0.7\text{V}$ $T_s=5760\text{K}$ $T_a=300\text{K}$ $P_s=1275\text{W/m}^2$

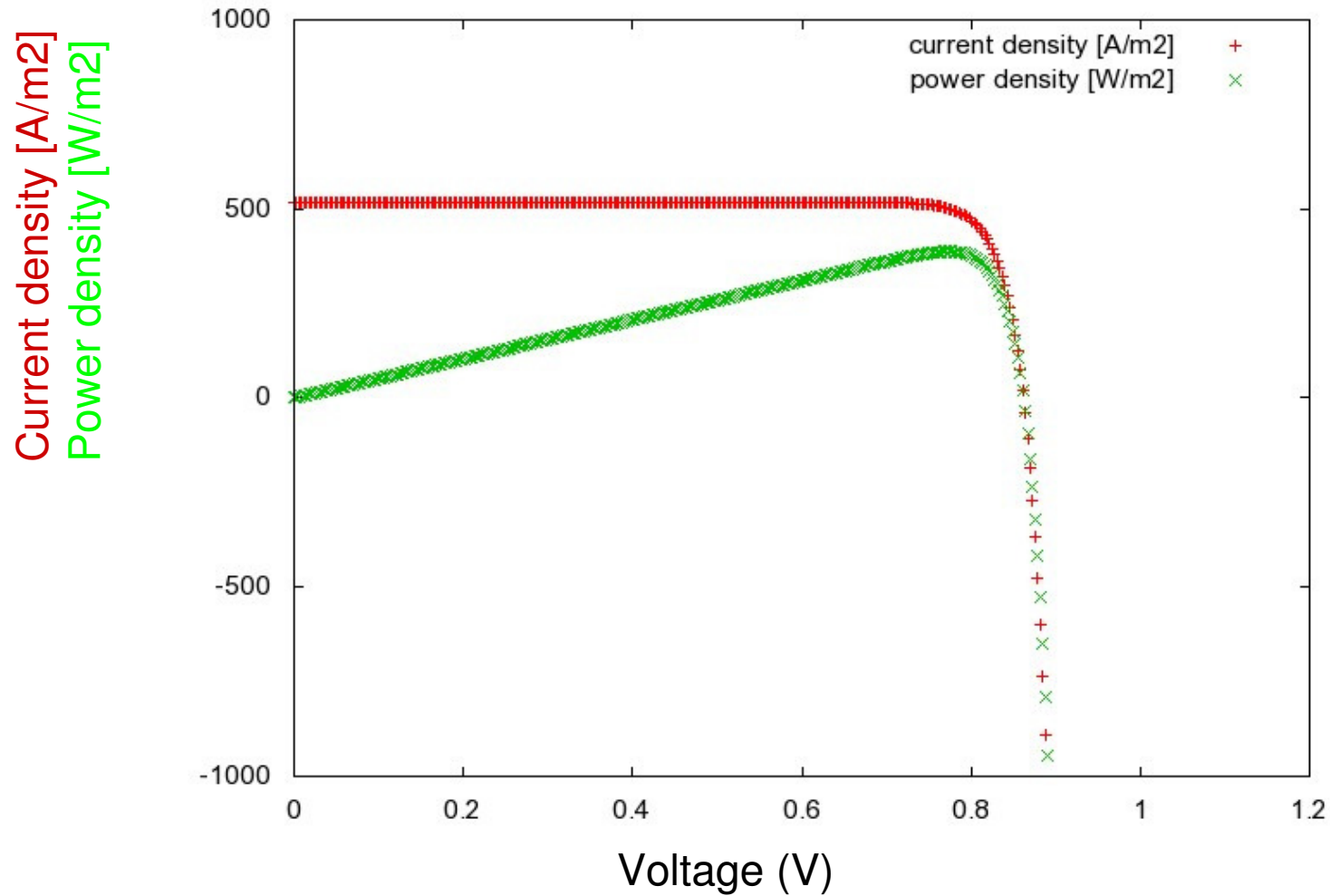
Maximum efficiency for $E_g=0.7\text{eV}$



$$\eta = \frac{VJ(V)}{P_s}$$

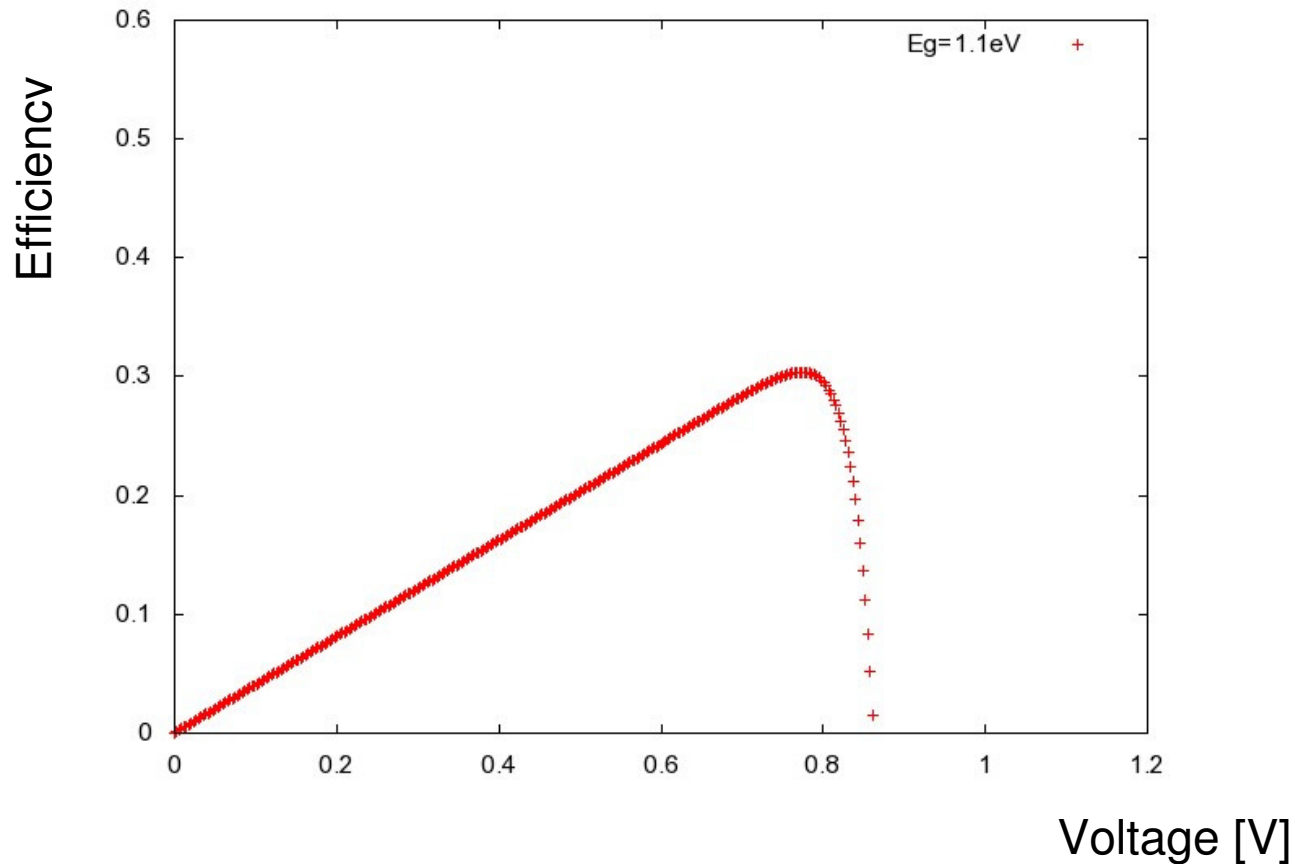
Best performance at $V_m=0.45\text{V}$ with 22.5%

Maximum current and power densities with $E_g=1.1\text{eV}$



$E_g=0.7\text{V}$ $T_s=5760\text{K}$ $T_a=300\text{K}$ $P_s=1275\text{W/m}^2$

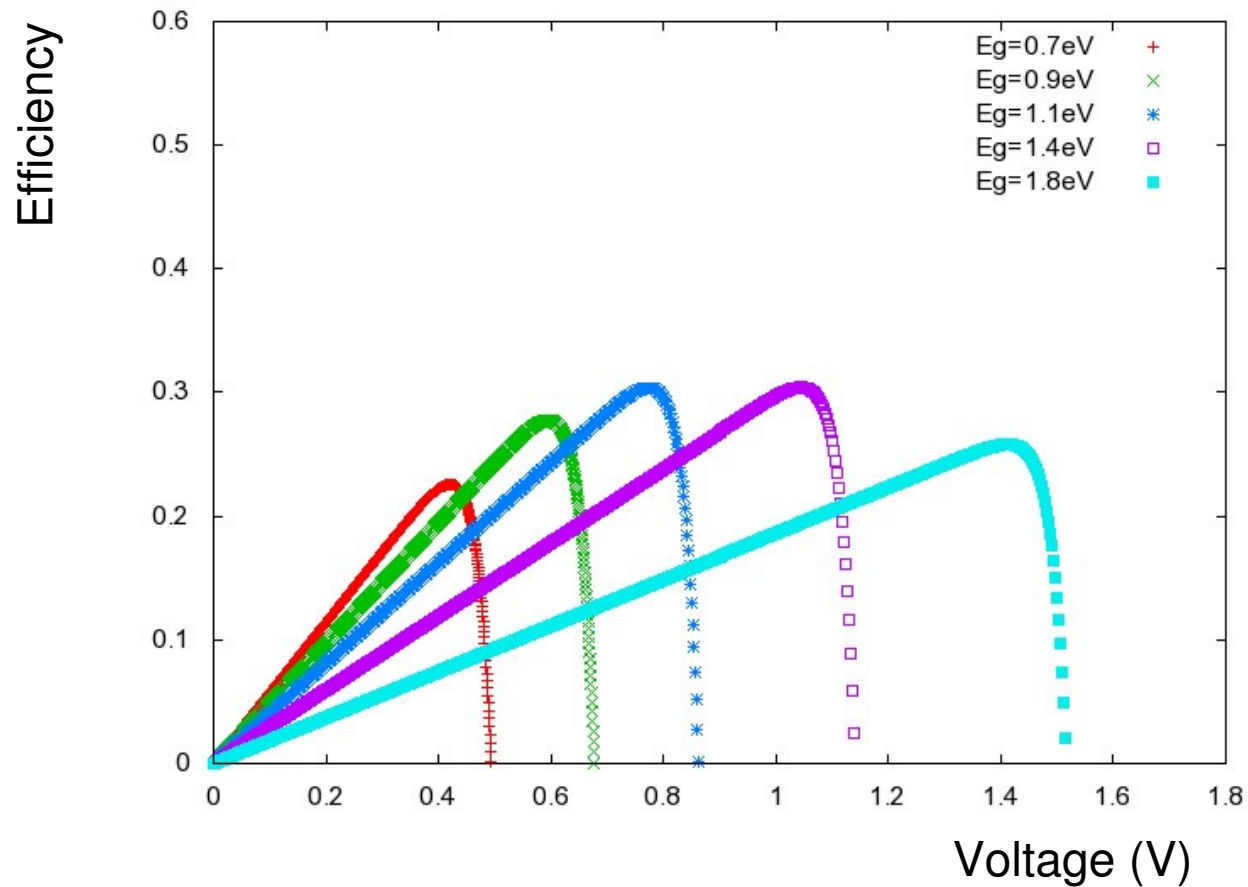
Maximum efficiency for $E_g=1.1\text{eV}$



$$\eta = \frac{VJ(V)}{P_s}$$

Best performance at $V_m=0.77\text{V}$ with 30%

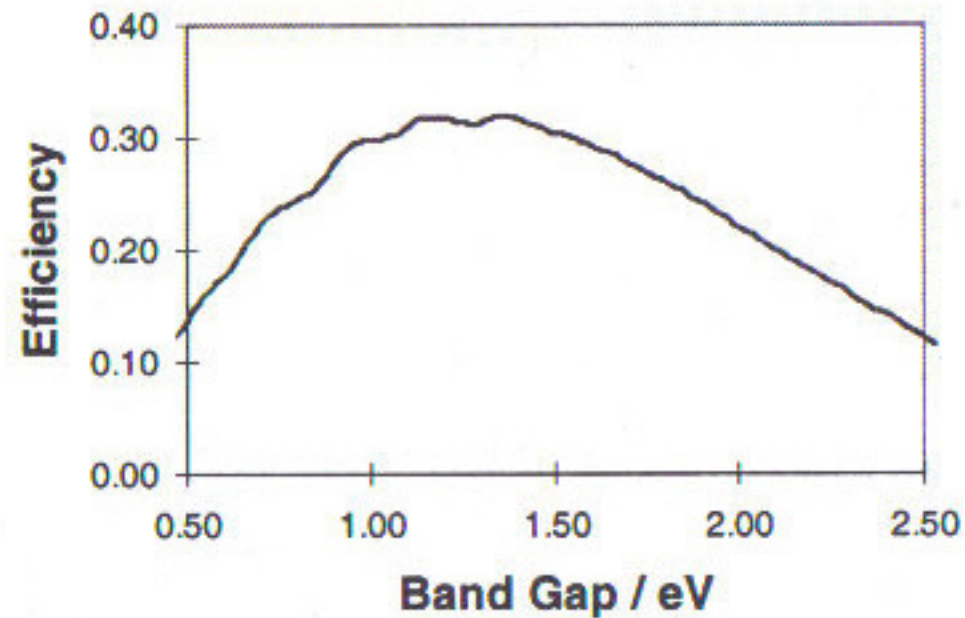
Maximum efficiency for different band gaps



$$\eta = \frac{VJ(V)}{P_s}$$

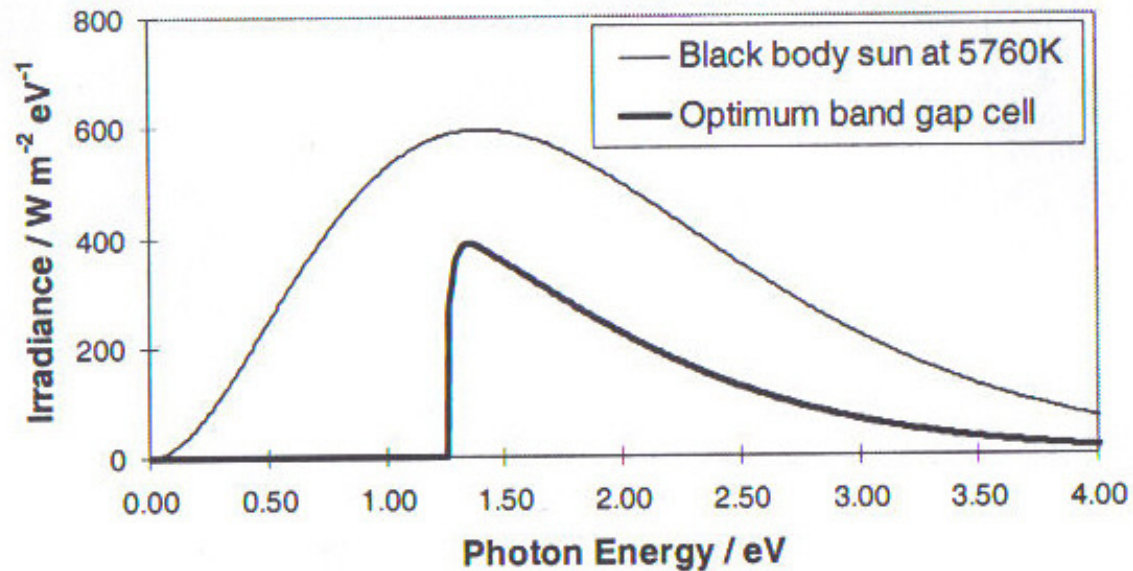
E_g=0.7-1.8V T_s=5760K T_a=300K P_s=1275W/m²

Limiting efficiency for a single band gap cell in AM 1.5



Optimum of a single band gap cell is 33% at 1.4eV

Power spectrum of a black body at 5760K and power available to a optimum band gap cell



Photons with $E < E_g$ are not used

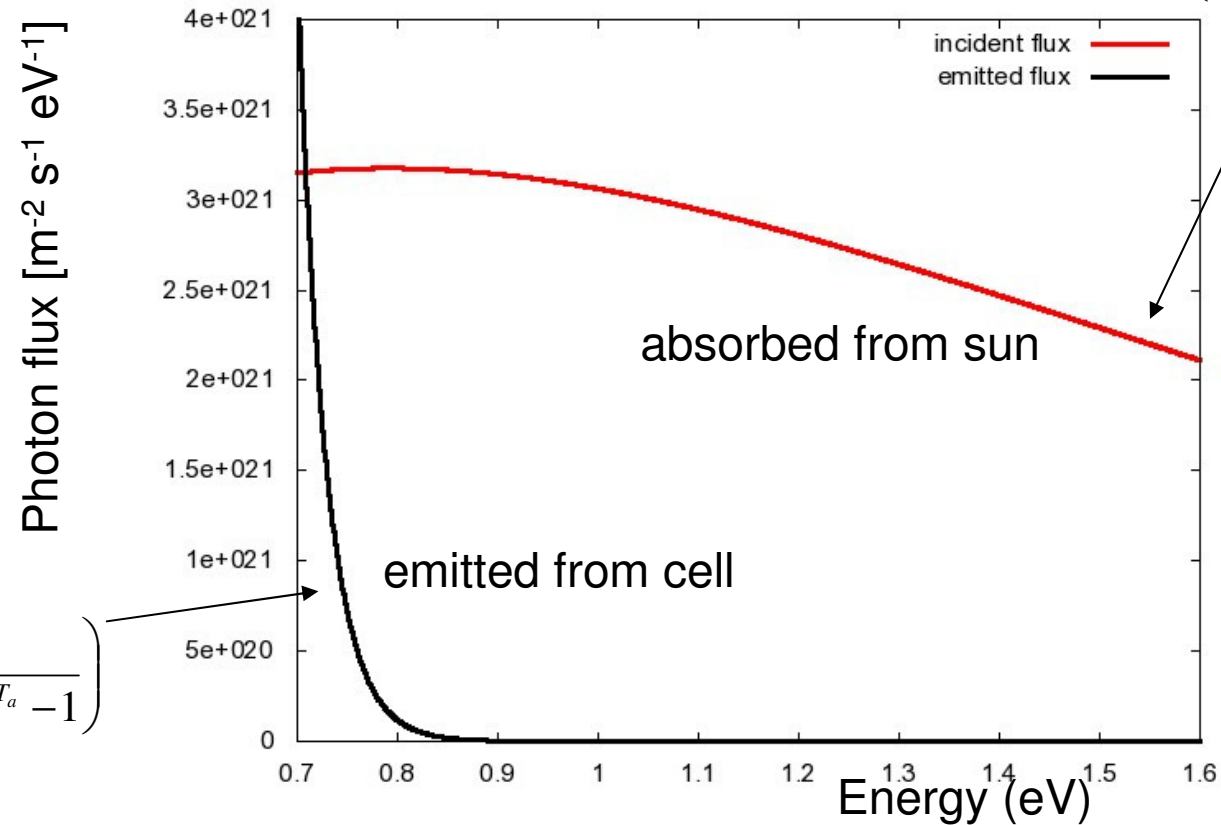
Photons with $E \gg E_g$ have the same effect as $E = E_g$

Comparison of Photon Fluxes

$(V < V_{oc})$

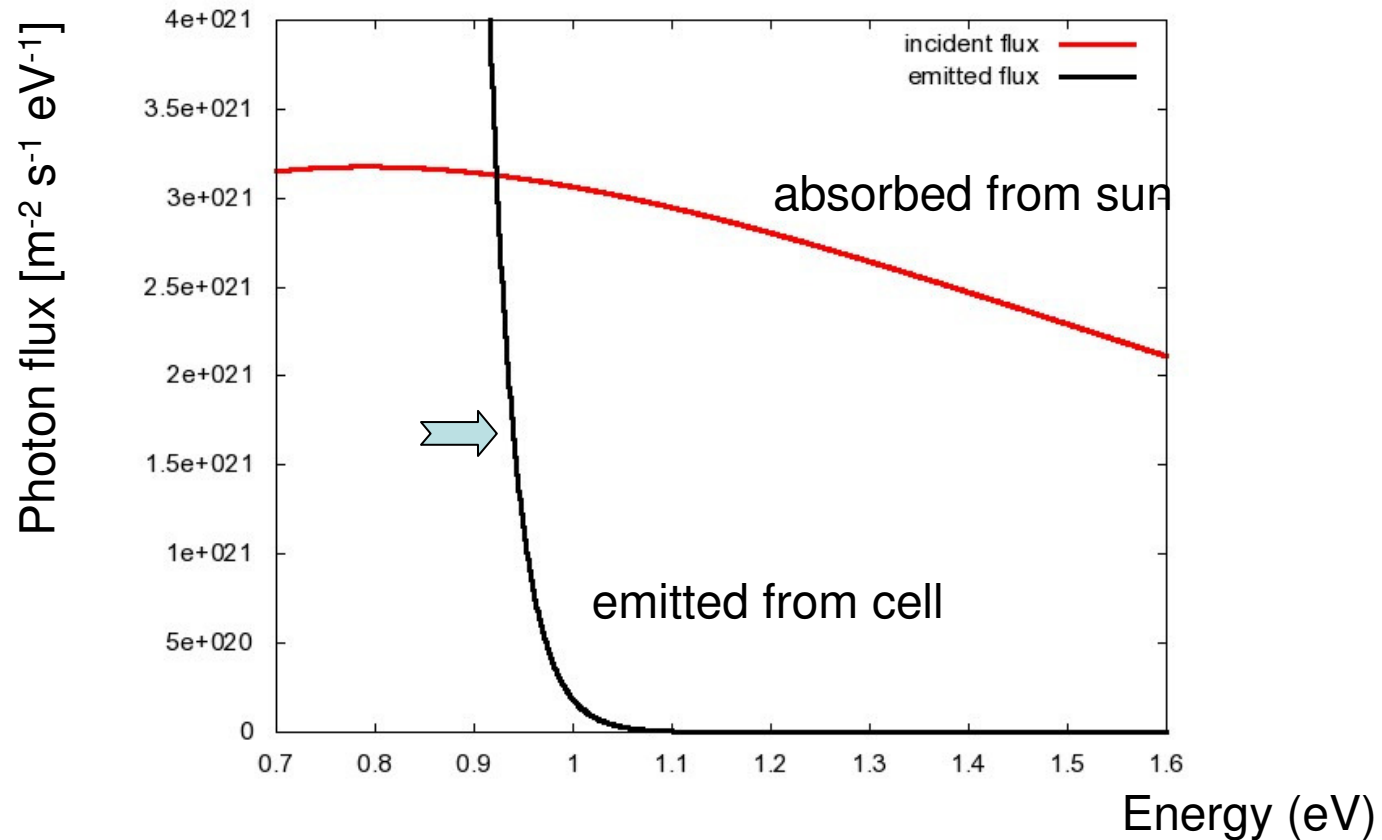
$$b_s(E) = \frac{2F_s}{h^3 c^2} \left(\frac{E^2}{E^{E/k_B T_s} - 1} \right)$$

$$b_e(E) = \frac{2F_a}{h^3 c^2} \left(\frac{E^2}{E^{(E-qV)/k_B T_a} - 1} \right)$$



$V=0.4V$ $E_g=0.7V$ $T_s=5760K$ $T_a=300K$ $P_S=1275W/m^2$

Comparison of Photon Fluxes ($V > V_{oc}$)



$$V=0.6\text{V} \quad E_g=0.7\text{V} \quad T_s=5760\text{K} \quad T_a=300\text{K} \quad P_S=1275\text{W/m}^2$$

Requirements for ideal photoconverter

- Energy gap E_g
- All incident radiation with $E > E_g$ is absorbed
- Each photon creates one electron-hole pair
- Excited pairs do not recombine, except radiatively, as required by the balance
- Charges are completely separated
- Charges are transported without losses

Energy gap E_g

The optimum band gap for AM1.5 of 1.4eV with 33%

III-V-semiconductors: GaAs 1.42eV (2009: 26% achieved)
InP 1.35eV

but expensive

Silicon 1.1eV would give a maximum efficiency of 30% (2009: 24.7%)
Cheap and high quality crystals available, but energy intensive production

CdTe and CuInGaSe_2 are developed and have low energy production costs
(2009: 19.9%)

Other parameters

Light absorption: Almost perfect absorption with $E > E_g$ is achieved for layer thicknesses of 10-100 microns

Charge separation: Asymmetry is realized with pn-junctions (either highly doped layers of same materials or different materials)
Junction quality is of central importance. (especially large areas)

Lossless transport:

Good Ohmic contacts, low resistance material for conduction (serial resistance),
No leakage current (parallel resistance)

Optimum load resistance:

Should be optimized to match best operating point

For most applications rather arrays than individual cells are to be optimized

Reasons for losses

- Reflection at front surface
- Non-radiative recombination of carriers
(defect sites trap charges, which subsequently recombine)
- Voltage drop due to series resistance $\Delta\mu \neq qV$

| Classification ^a | Effic. ^b (%) | Area ^c (cm ²) | V _{oc} (V) | J _{sc} (mA/cm ²) | FF ^d (%) | Test centre ^e (and date) | Description |
|----------------------------------|----------------------------|---|------------------------|--|------------------------|--|---|
| Silicon | | | | | | | |
| Si (crystalline) | 25.0 ± 0.5 | 4.00 (da) | 0.705 | 42.7 | 82.8 | Sandia (3/99) ^f | UNSW PERL ¹² |
| Si (multicrystalline) | 20.4 ± 0.5 | 1.002 (ap) | 0.664 | 38.0 | 80.9 | NREL (5/04) ^f | FhG-ISE ¹³ |
| Si (thin film transfer) | 16.7 ± 0.4 | 4.017 (ap) | 0.645 | 33.0 | 78.2 | FhG-ISE (7/01) ^f | U. Stuttgart (45 μm thick) ¹⁴ |
| Si (thin film submodule) | 10.5 ± 0.3 | 94.0 (ap) | 0.492 ^g | 29.7 ^g | 72.1 | FhG-ISE (8/07) ^f | CSG Solar (1-2 μm on glass; 20 cells) ¹⁵ |
| III-V cells | | | | | | | |
| GaAs (crystalline) | 26.1 ± 0.8 | 0.998 (ap) | 1.038 | 29.7 | 84.7 | FhG-ISE (12/07) ^f | Radboud U. Nijmegen ⁶ |
| GaAs (thin film) | 26.1 ± 0.8 | 1.001 (ap) | 1.045 | 29.5 | 84.6 | FhG-ISE (07/08)^f | Radboud U. Nijmegen⁶ |
| GaAs (multicrystalline) | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| InP (crystalline) | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| Thin film chalcogenides | | | | | | | |
| CIGS (cell) | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| CIGS (submodule) | 10.5 ± 0.3 | 94.0 (ap) | 0.492 ^g | 29.7 ^g | 72.1 | FhG-ISE (8/07) ^f | CSG Solar (1-2 μm on glass; 20 cells) ¹⁵ |
| CdTe (cell) | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| Amorphous/nanocrystalline | | | | | | | |
| Si (amorphous) | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| Si (nanocrystalline) | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| Photochemical | | | | | | | |
| Dye sensitised | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| Dye sensitised (series) | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| Dye sensitised (submodule) | 10.5 ± 0.3 | 94.0 (ap) | 0.492 ^g | 29.7 ^g | 72.1 | FhG-ISE (8/07) ^f | CSG Solar (1-2 μm on glass; 20 cells) ¹⁵ |
| Organic | | | | | | | |
| Organic polymer | 18.4 ± 0.5 | 4.011 (ap) | 0.604 | 33.3 | 79.7 | NREL (11/05) ^f | PTL Ga substrate ¹⁶ |
| Organic (submodule) | 10.5 ± 0.3 | 94.0 (ap) | 0.492 ^g | 29.7 ^g | 72.1 | FhG-ISE (8/07) ^f | CSG Solar (1-2 μm on glass; 20 cells) ¹⁵ |
| Multijunction devices | | | | | | | |
| GaInP/GaAs/Ge | 32.0 ± 1.5 ^j | 3.989 (t) | 2.622 | 14.37 | 85.0 | NREL (1/03) | Spectrolab (monolithic) |
| GaInP/GaAs | 30.3 ^j | 4.0 (t) | 2.488 | 14.22 | 85.6 | JQA (4/96) | Japan Energy (monolithic) ²⁸ |

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Research

SHORT COMMUNICATION

Solar Cell Efficiency Tables (Version 33)

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on glass¹⁸

serial cells¹⁹

on glass²⁰

on glass)²²

l cells²⁴

cells²⁵

3HT/PCBM)²⁷

Ways to improve

- Concentrators
- Tandem cells
- Organic cells with a variety of band gaps

Types of solar cells with regard to technology

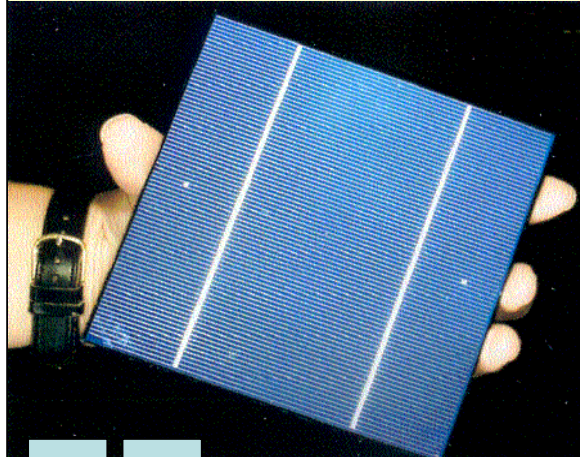
Concentrator systems

Ge / $\text{In}_x\text{Ga}_{1-x}\text{As}$ / $\text{In}_x\text{Ga}_{1-x}\text{P}$
world record: **41.6%**
module efficiency: 28%
system efficiency: 23%
direct sun light is demanded
need for optical tracking



Si-wafer-based

c-Si, mc-Si
world record : 24.7%
module efficiency : 18%
energy intensive
production of Si
(very high purity needed)
dominant on market



Thin film

a-Si:H, nc-Si, CdTe,
 $\text{Cu}(\text{In,Ga})(\text{S,Se})_2$
nanostructured solar cells
organic solar cells
world record :
19.9% $\text{Cu}(\text{In,Ga})\text{Se}_2$
integrated series
connection



Material mit Bandgaps

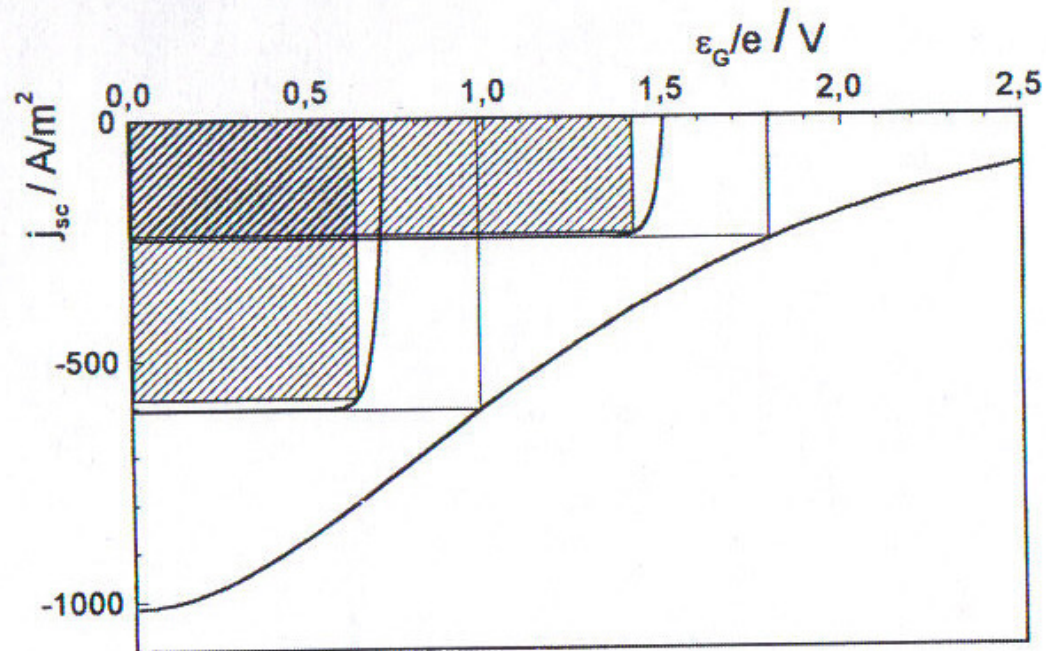


Abb. 8.2 Kennlinien von zwei Solarzellen mit Bandabstand $\epsilon_{G1} = 1.8 \text{ eV}$ und $\epsilon_{G2} = 0.98 \text{ eV}$

Theoretische Effizienz von Tandemzellen

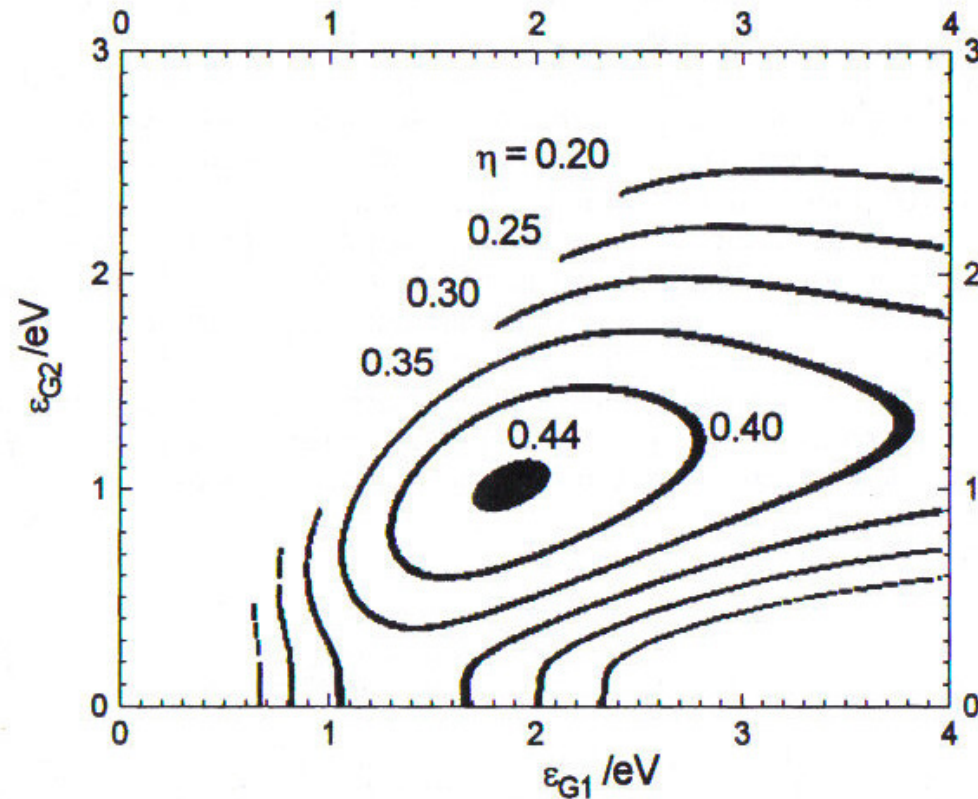
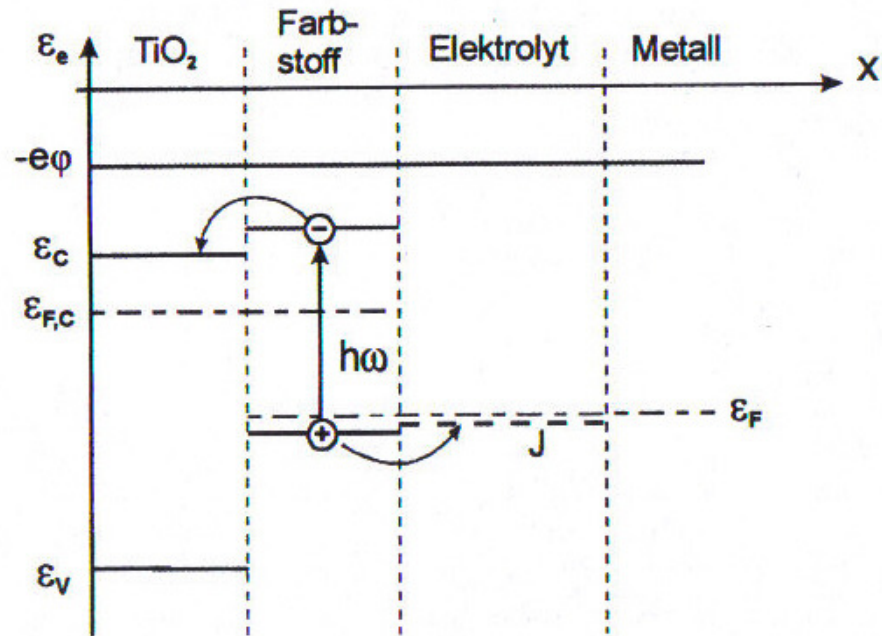


Abb. 8.3 Wirkungsgrad für zwei Tandemsolarzellen mit den Bandabständen ϵ_{G1} und ϵ_{G2} bei Addition ihrer Energieströme für das *AM0*-Spektrum

DSSC



Farbstoffsolarzelle, in der die Elektron-Loch Paare in dem Farbstoff Rutheniumbipyridil erzeugt werden. Die Elektronen fließen nach links über den n-Leiter TiO_2 ab, die Löcher nach rechts über Jod-Ionen, mit denen der Elektrolyt Acetonitril dotiert ist.

Farbstoff kann an Sonnenspektrum angepasst werden \Rightarrow verbesserter Wirkungsgrad
Probleme: Transport der Elektronen und Löcher zu den Elektroden \Rightarrow dünne Schicht
Absorption durch monomolekulare Schicht ist gering \Rightarrow Dickere Schichten